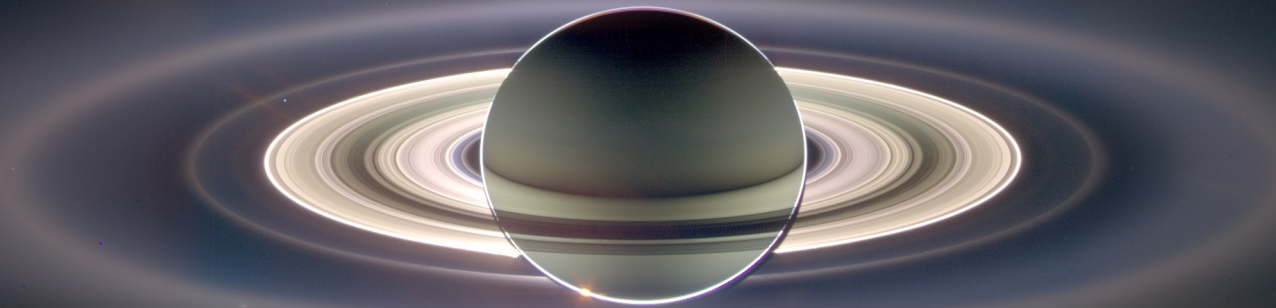
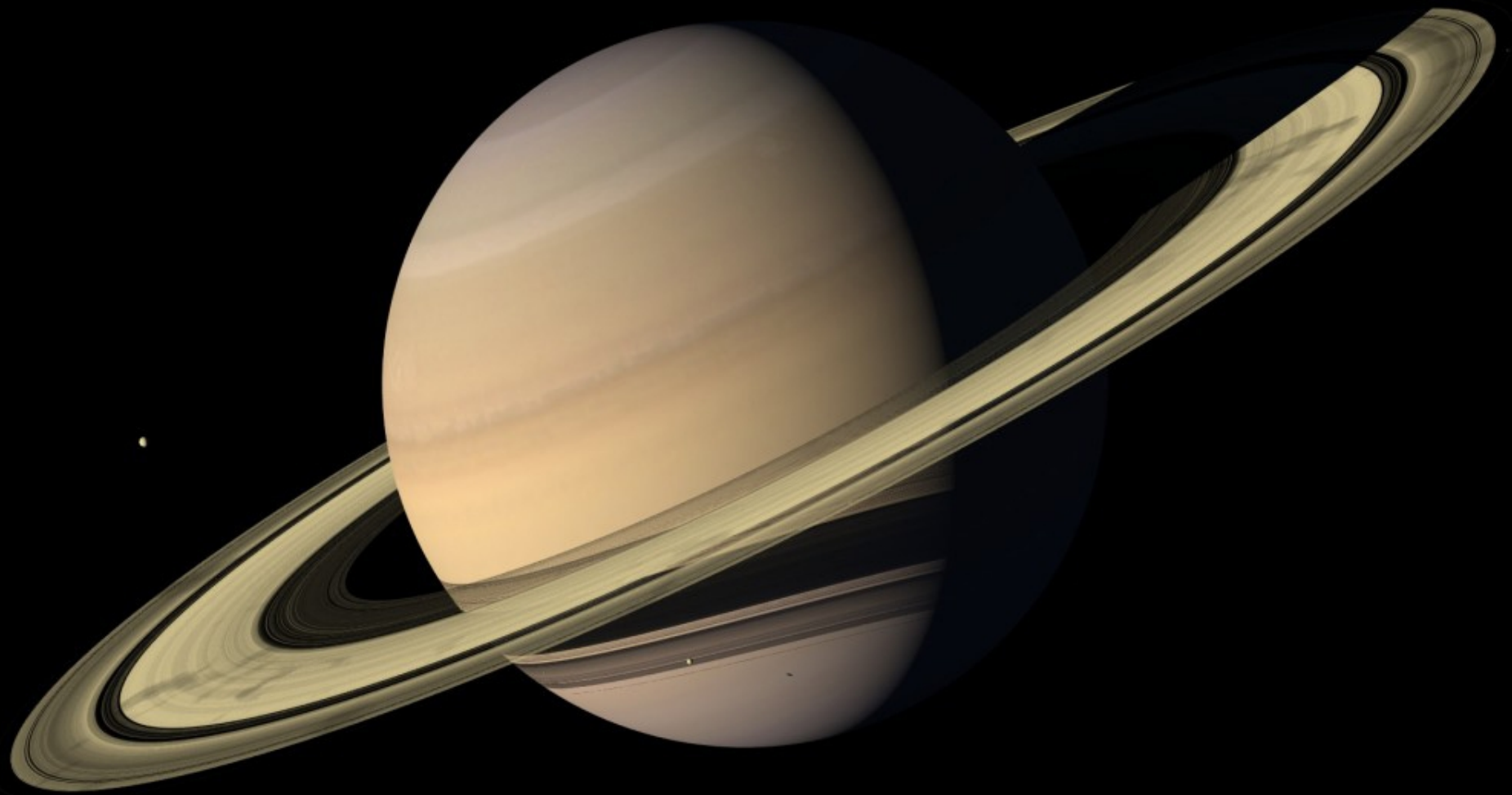


FORMATION OF SATELLITES from massive rings in the Solar system

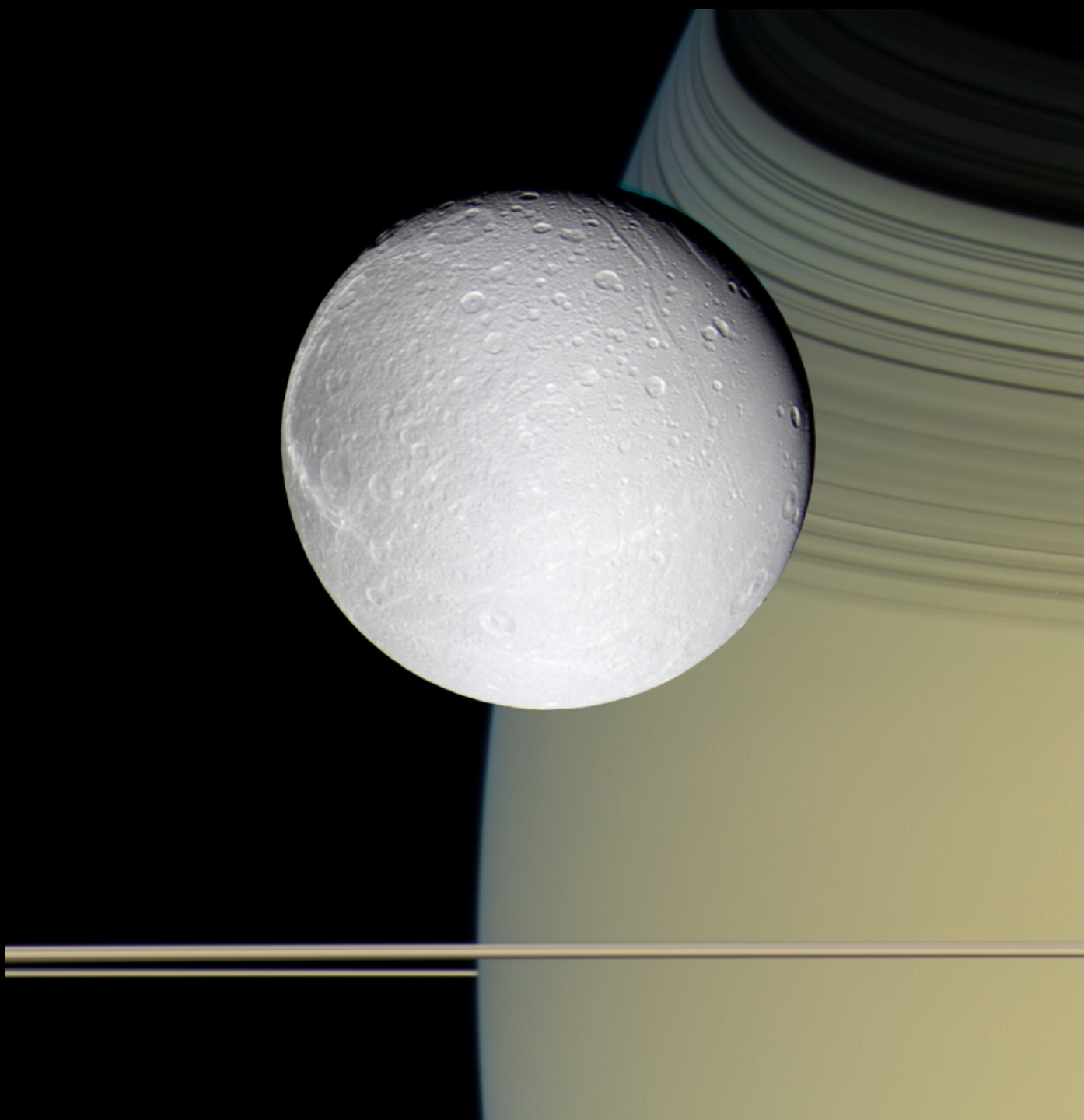


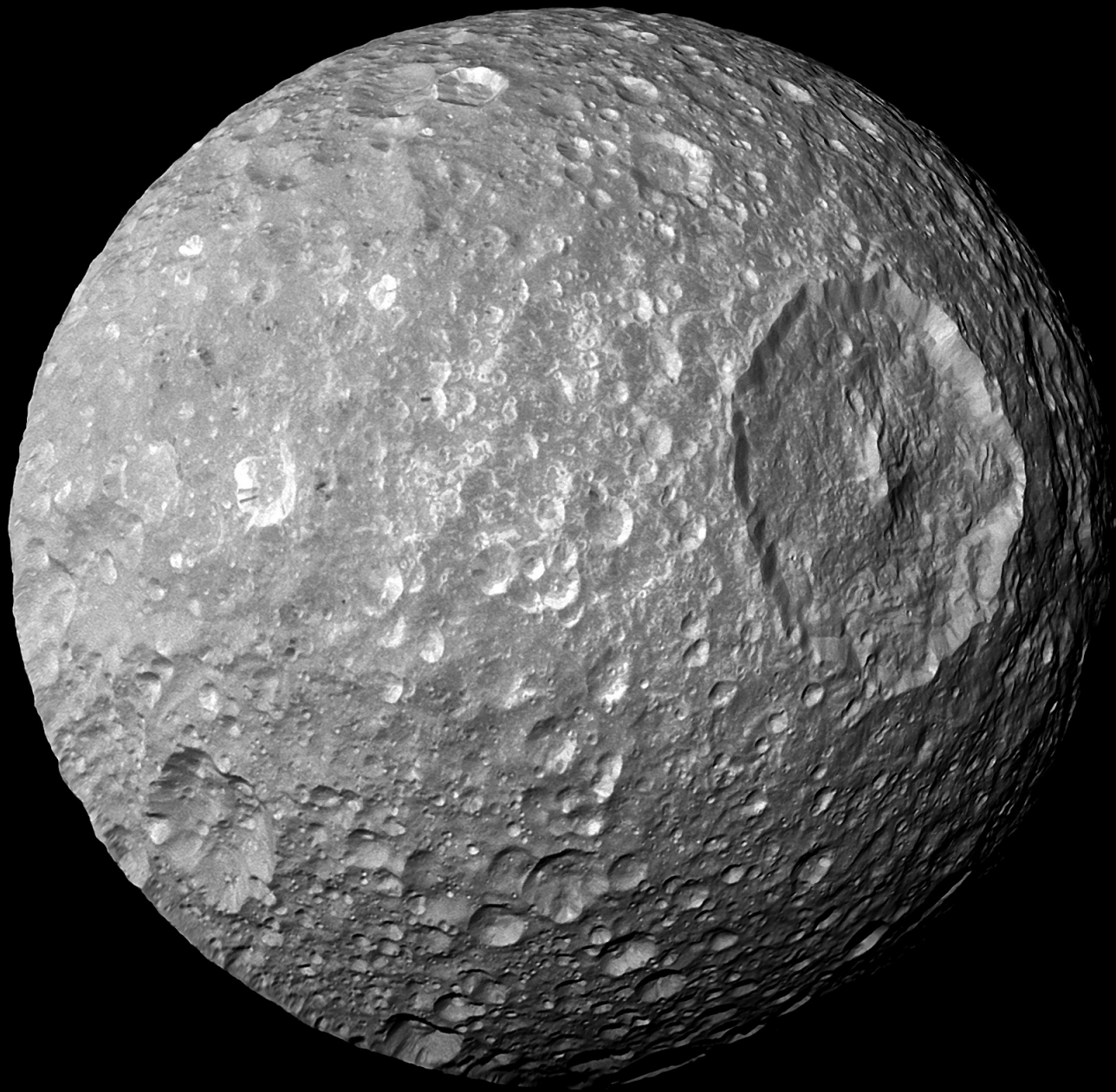
Aurélien CRIDA

& Sébastien CHARNOZ

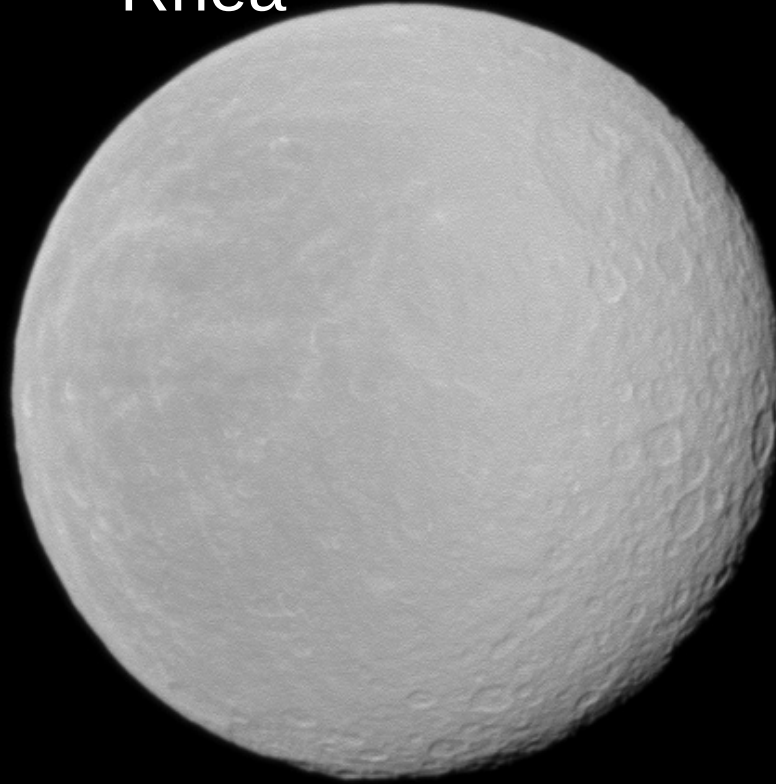








Rh ea



 pim eth e

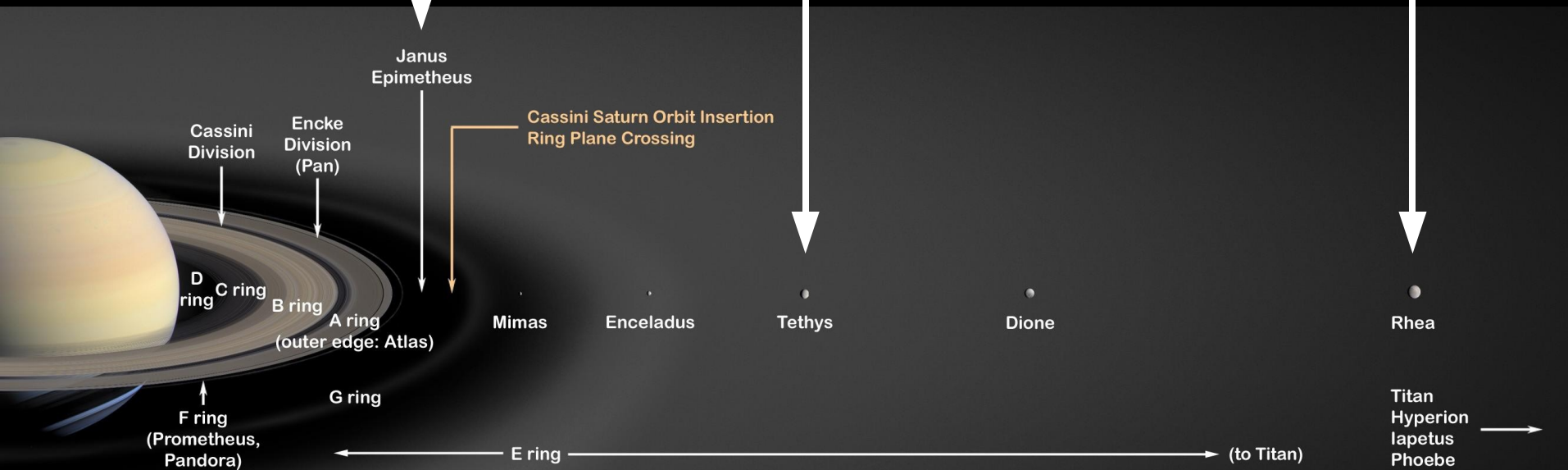
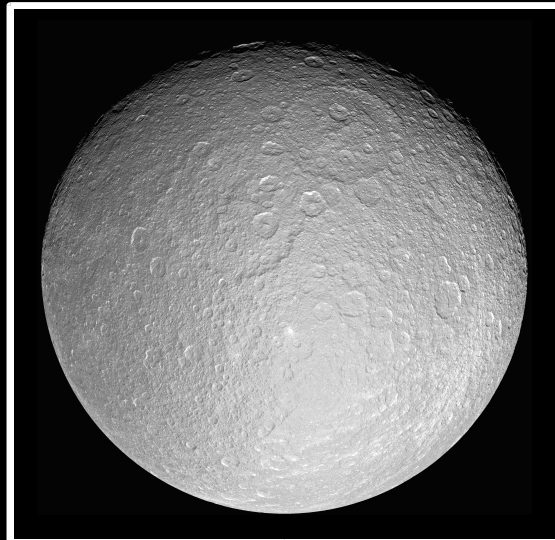


Dione

Prom eth e

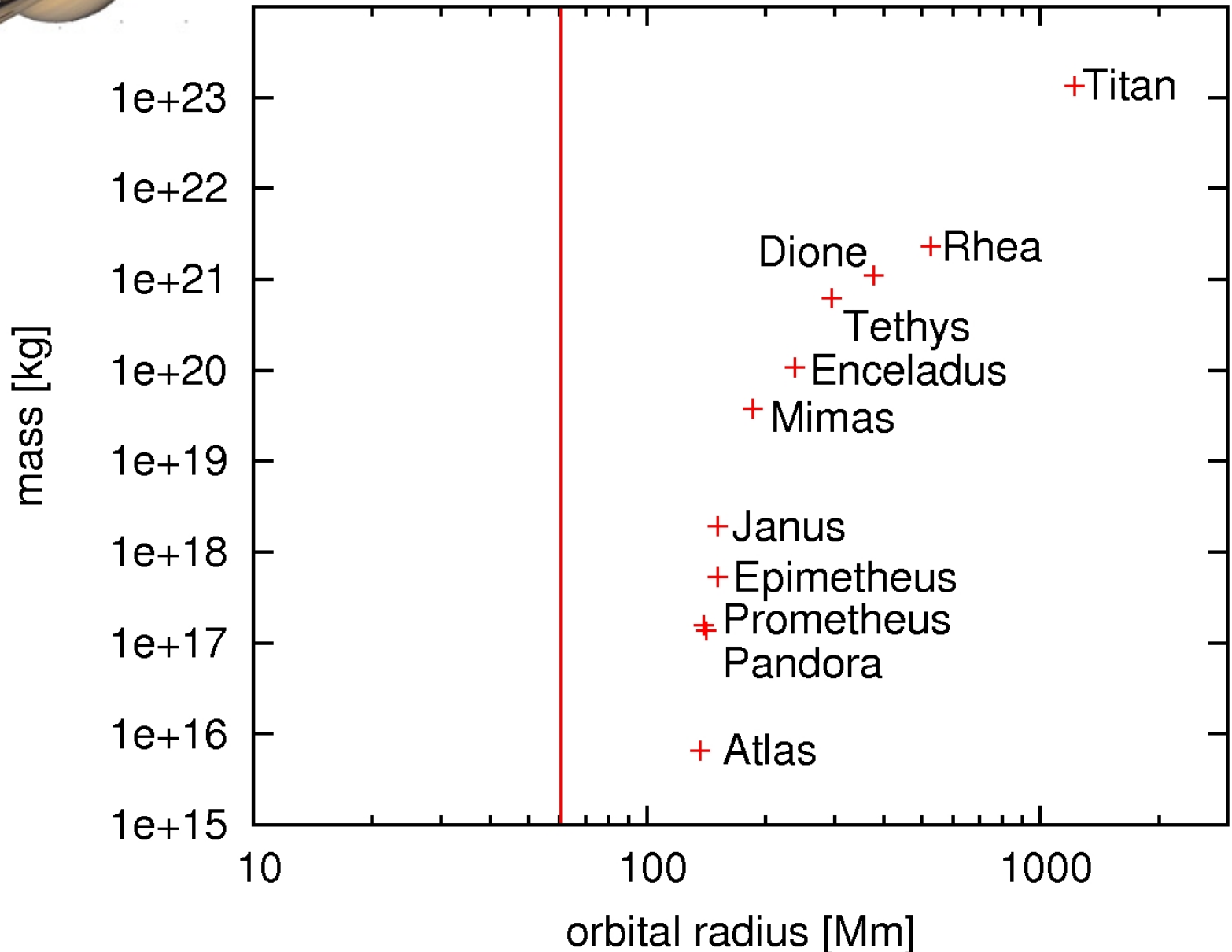
T thys





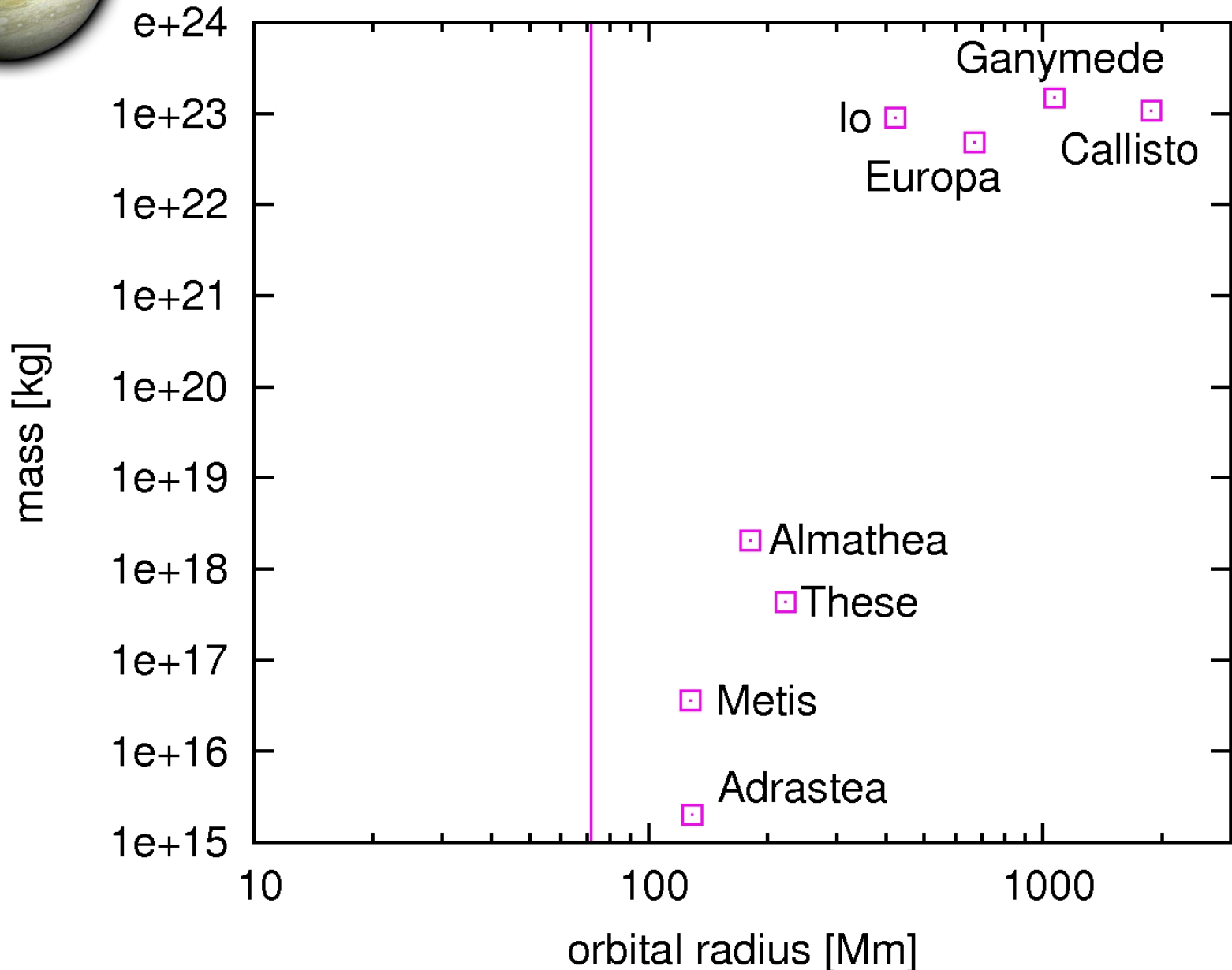


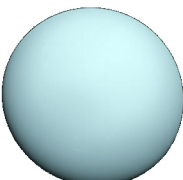
SATURN



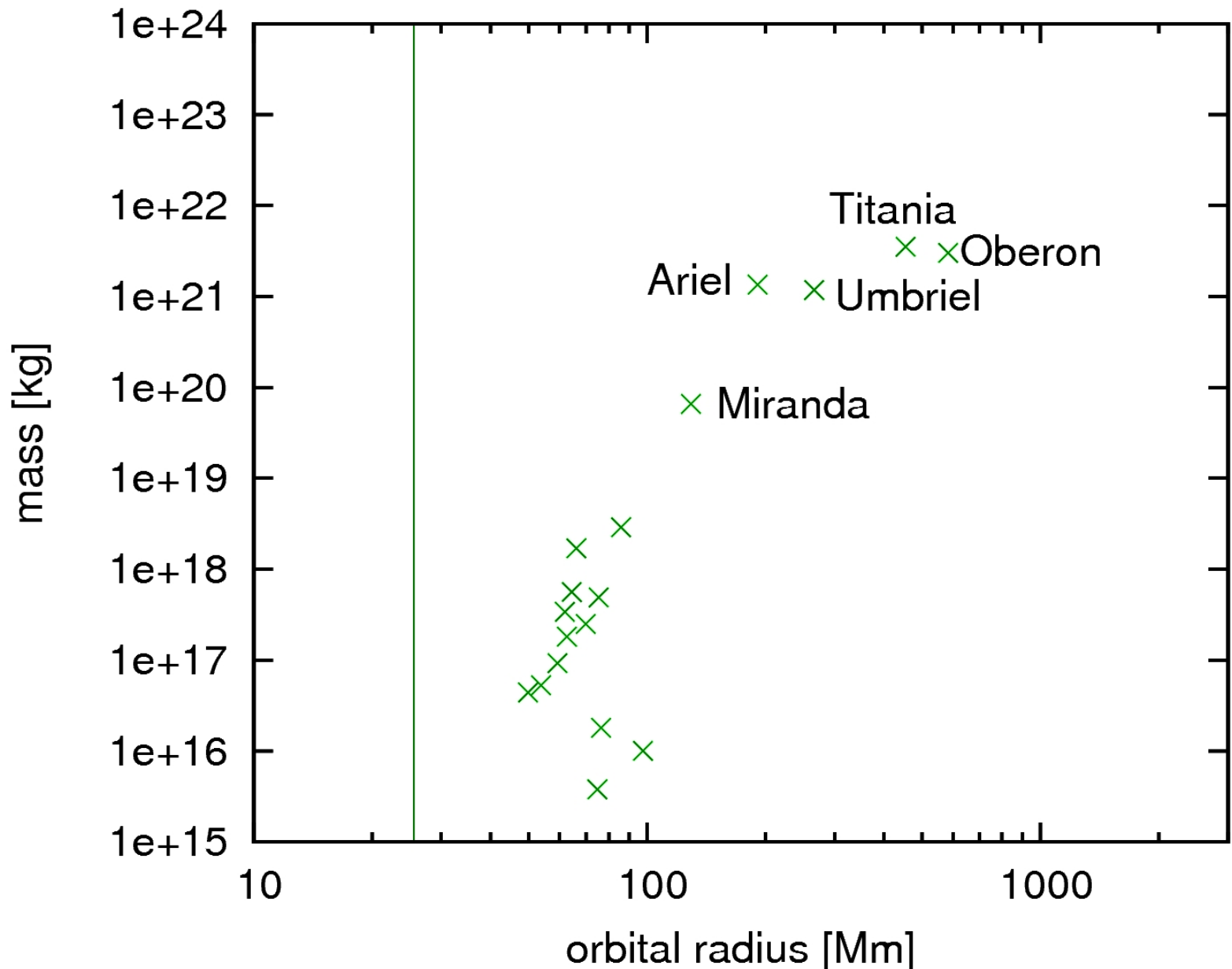


JUPITER



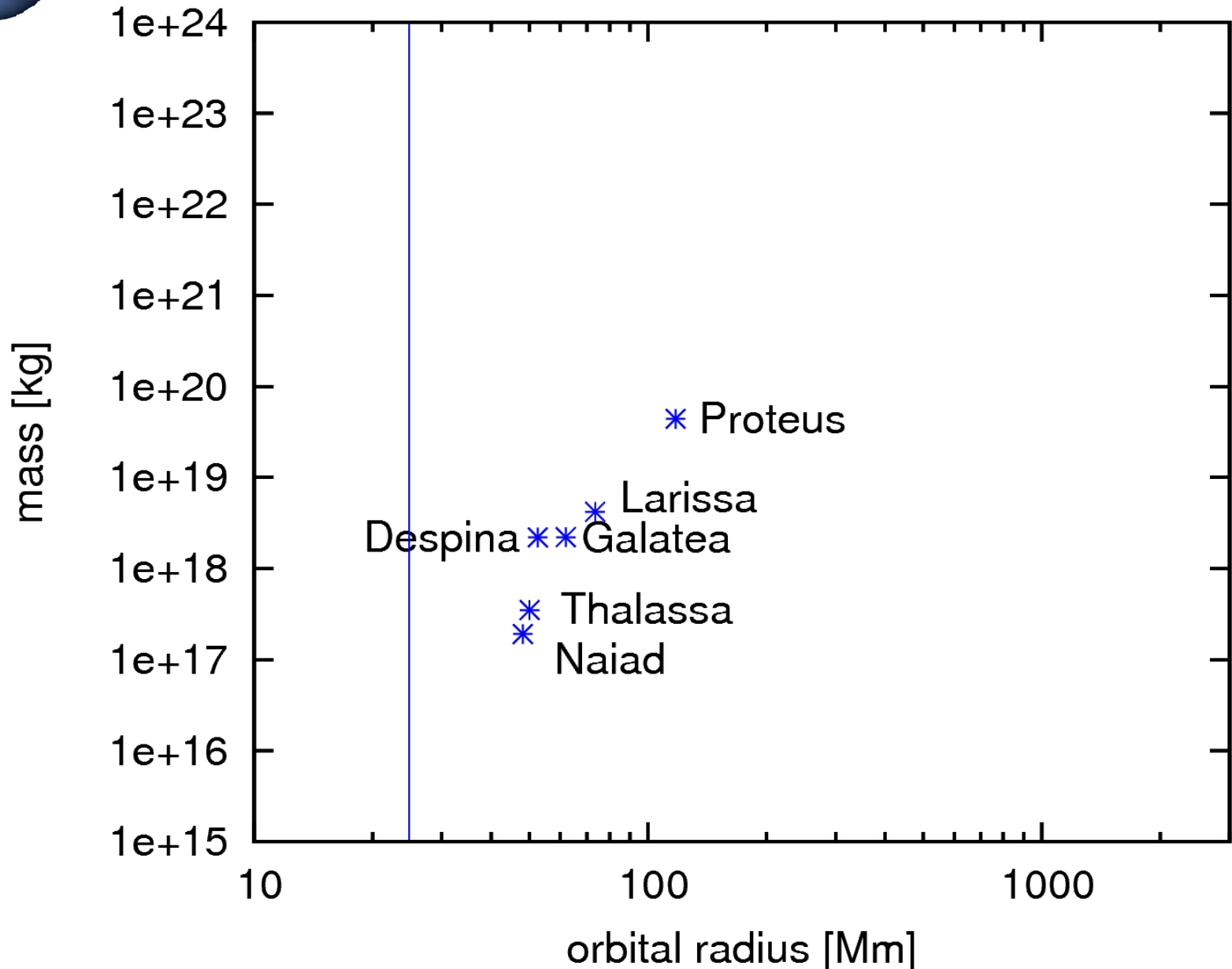


URANUS

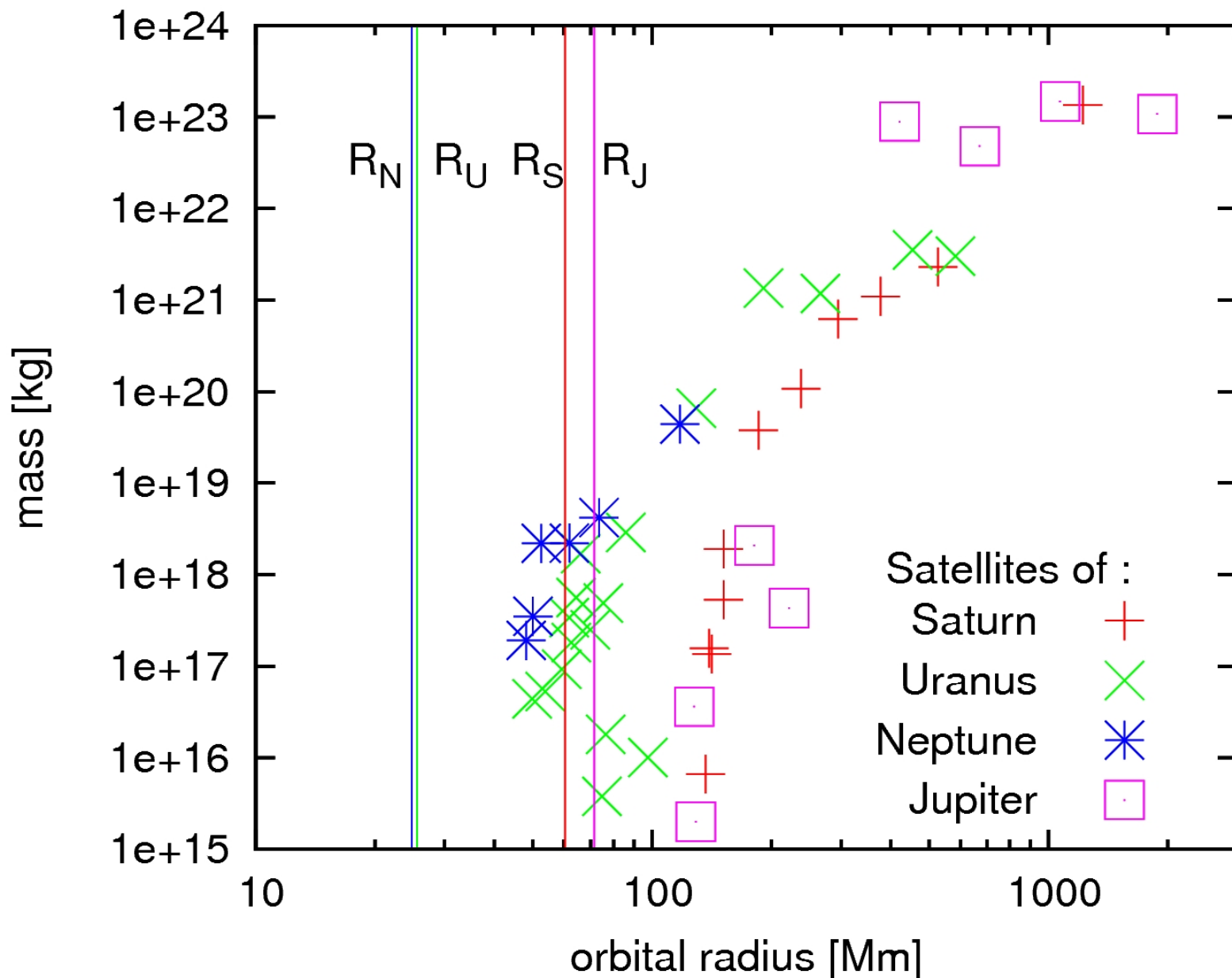




NEPTUNE



ALL GIANT PLANETS



ORIGIN of the SATELLITES ?

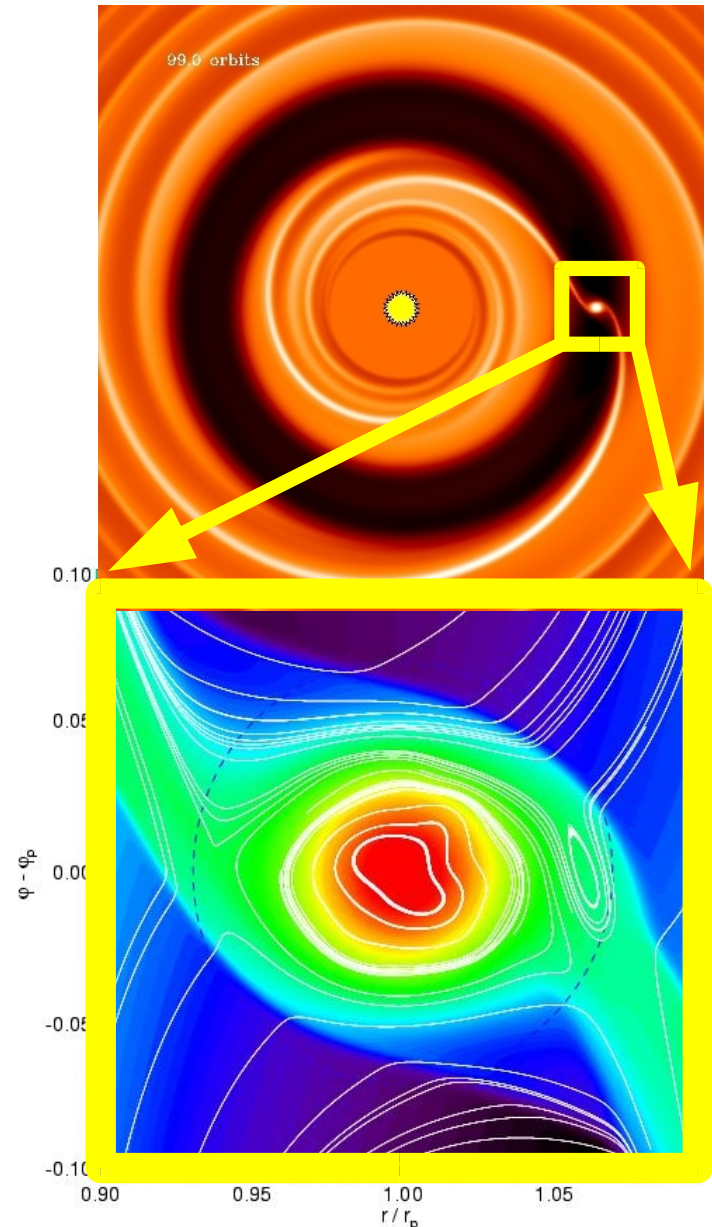
Planets form in a gas/dust disk around the Sun.

A giant planet carves a gap in the disk, and has its own circum-planetary disk.

A mini-planetary system would then form around the planet.

(Canup & Ward 2002, 2006 ;
Sasaki et al. 2010 ;
Mosqueira & Estrada 2003a,b...)

But the satellites systems don't look like the Solar System.

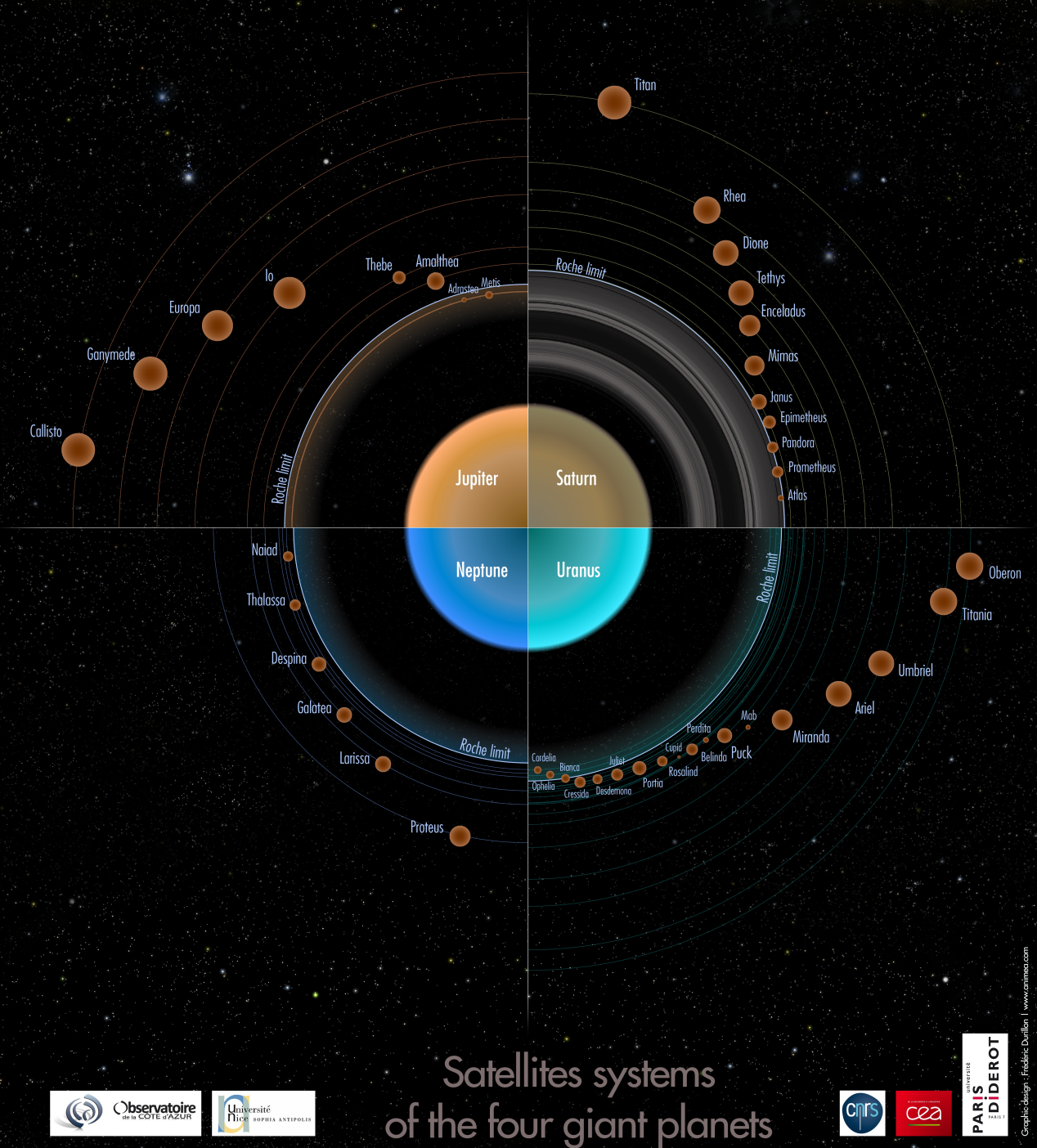


Distributions of giant planets' regular satellites :

- don't reach the planet
- ranked by mass
- pile-up at a few planetary radii (small bodies)

Why ?

It's not a power law, which question the Circum-Planetary Disc model...

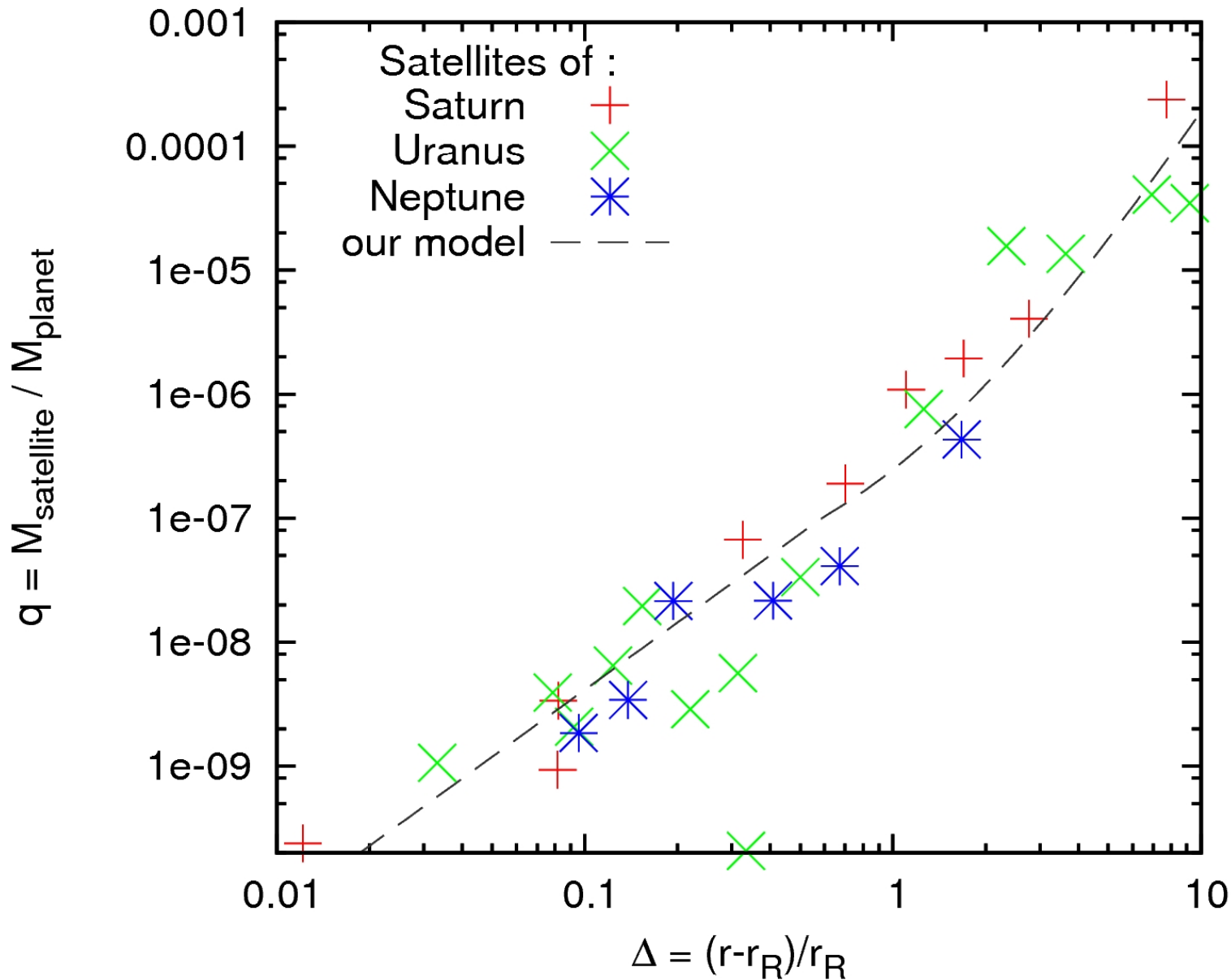


Satellites systems of the four giant planets



Graphic design: Frédéric Durillon | www.carnasa.com

CONCLUSION



TECHNICAL INTRODUCTION

1) Kepler's Law

2) Roche's radius

Kepler's Law

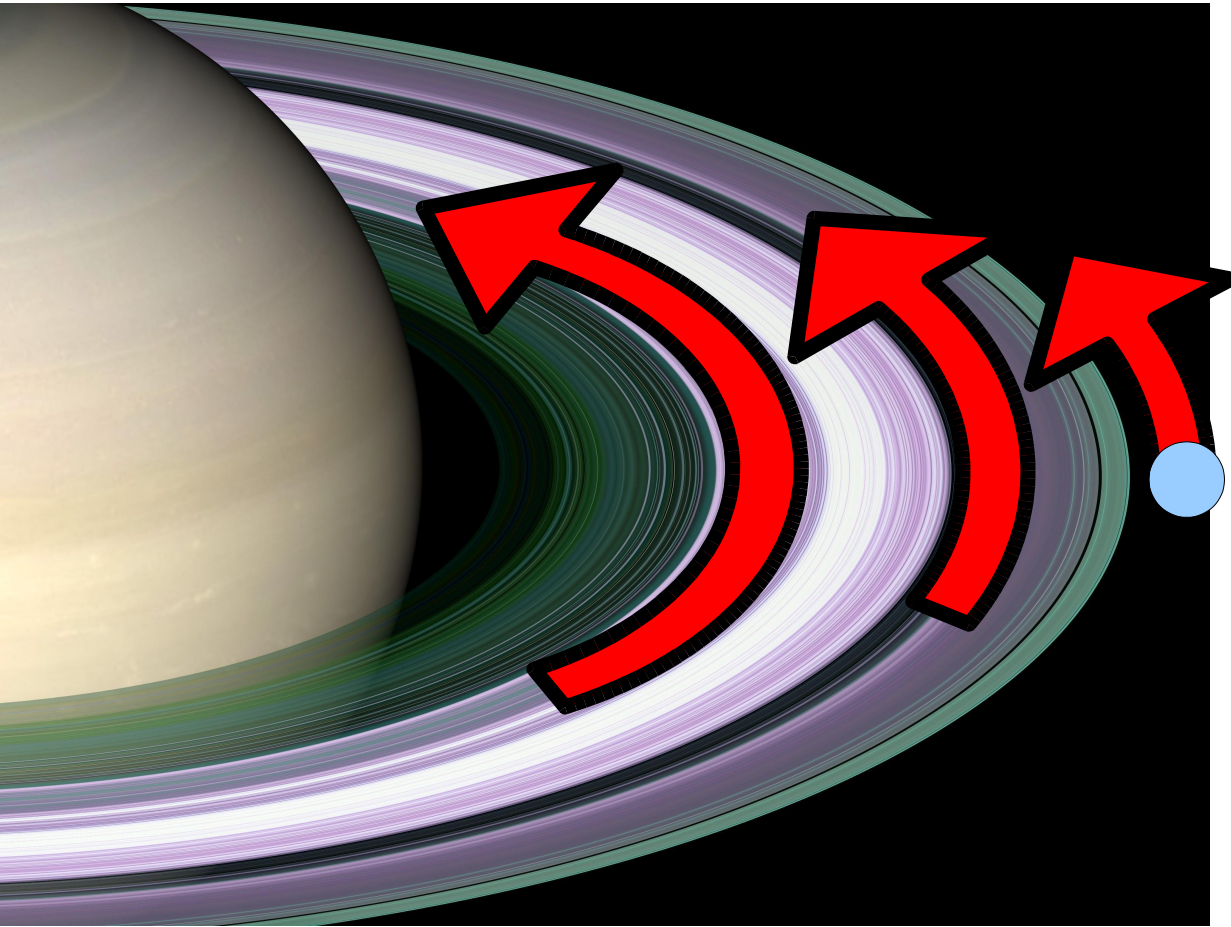


« The cube of the radius of an orbit is proportionnal to the square of the period. »

<u>planet:</u>	Mer	Earth	Jup
<u>period P [year]</u>	0,25	1	11,8
<u>radius R [AU]</u>	0,4	1	5,2
P^2 / R^3	1	1	1

OR : The closer from the Sun (or Saturn), the faster the rotation.

Kepler's Law : Application

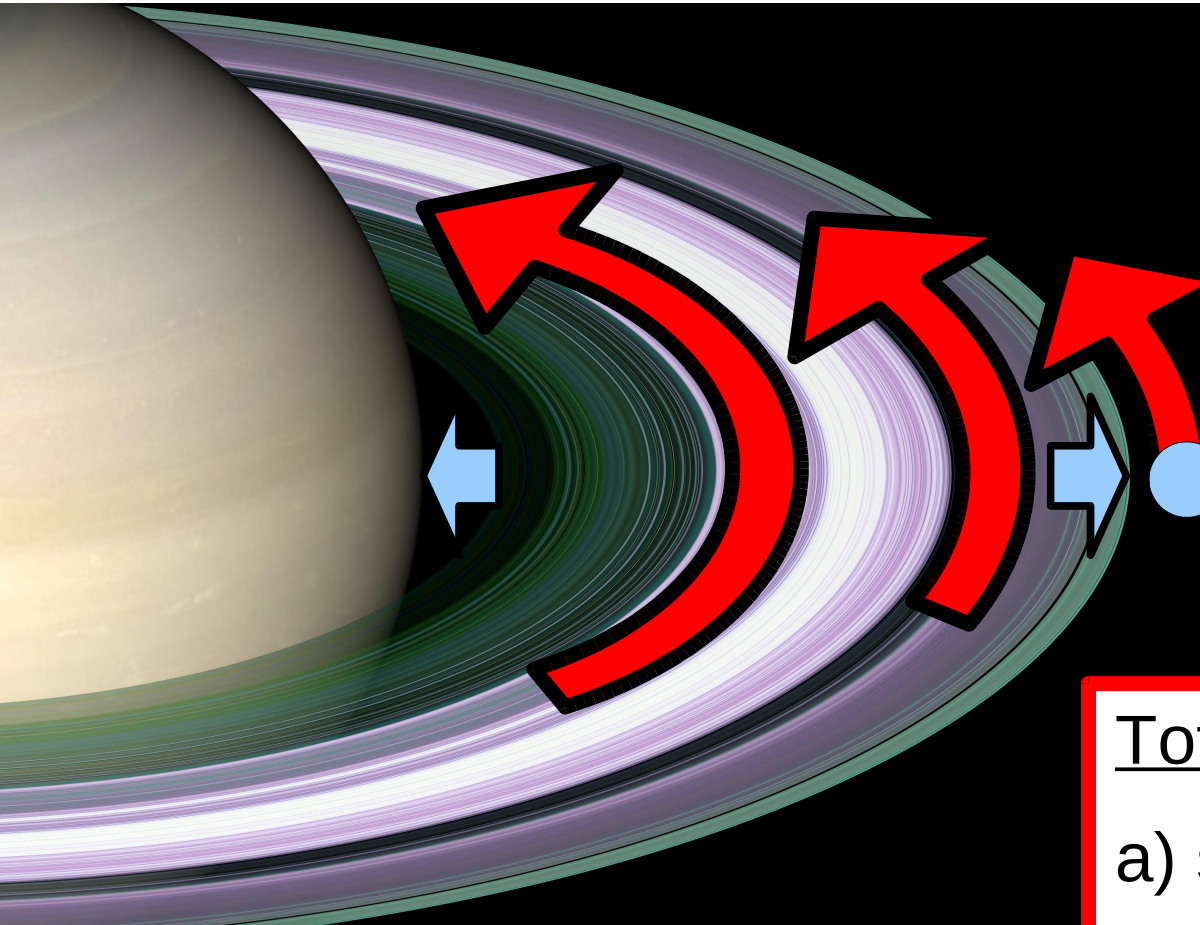


Take 2 bodies in orbit around Saturn (within the rings or outside).

The inner one rotates faster than the outer one.

But they attract each other, which accelerates the outer one, and slows down the inner one.

Kepler's Law : Application

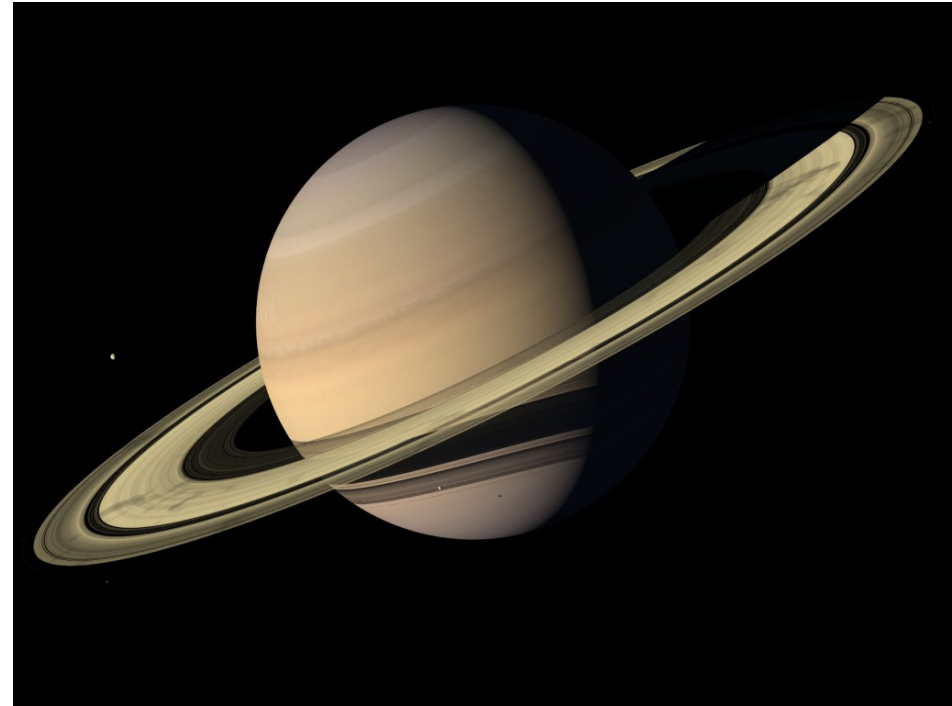
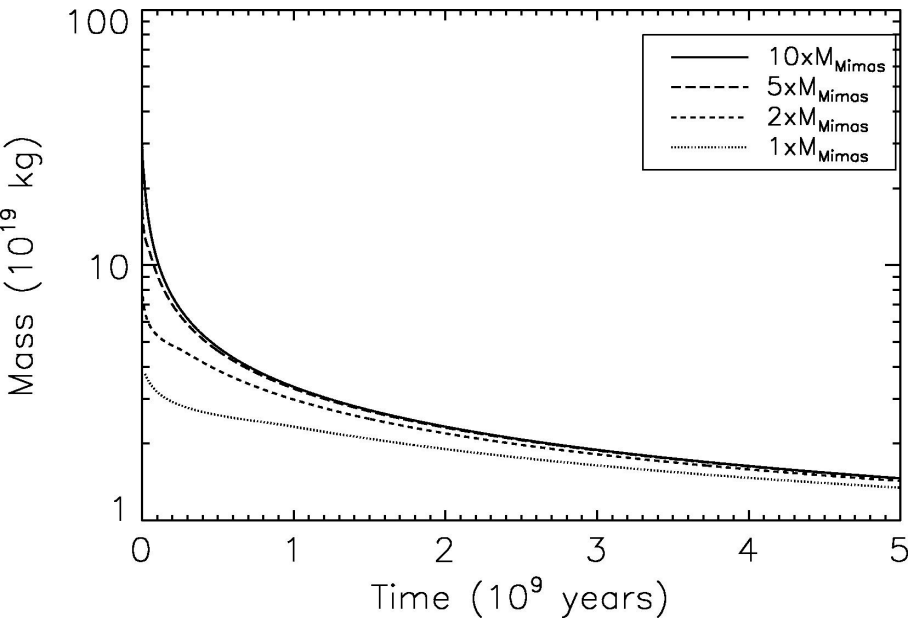


The outside is therefore thrown further, and the inside closer to the planet.

Total:

- a) spreading of the rings
- b) outward migration of the satellite.

Spreading of the rings



We found that independantly of the initial mass, the rings should have their present mass after 4.5 Gyrs.
(Salmon, Charnoz, Crida, Brahic 2010)

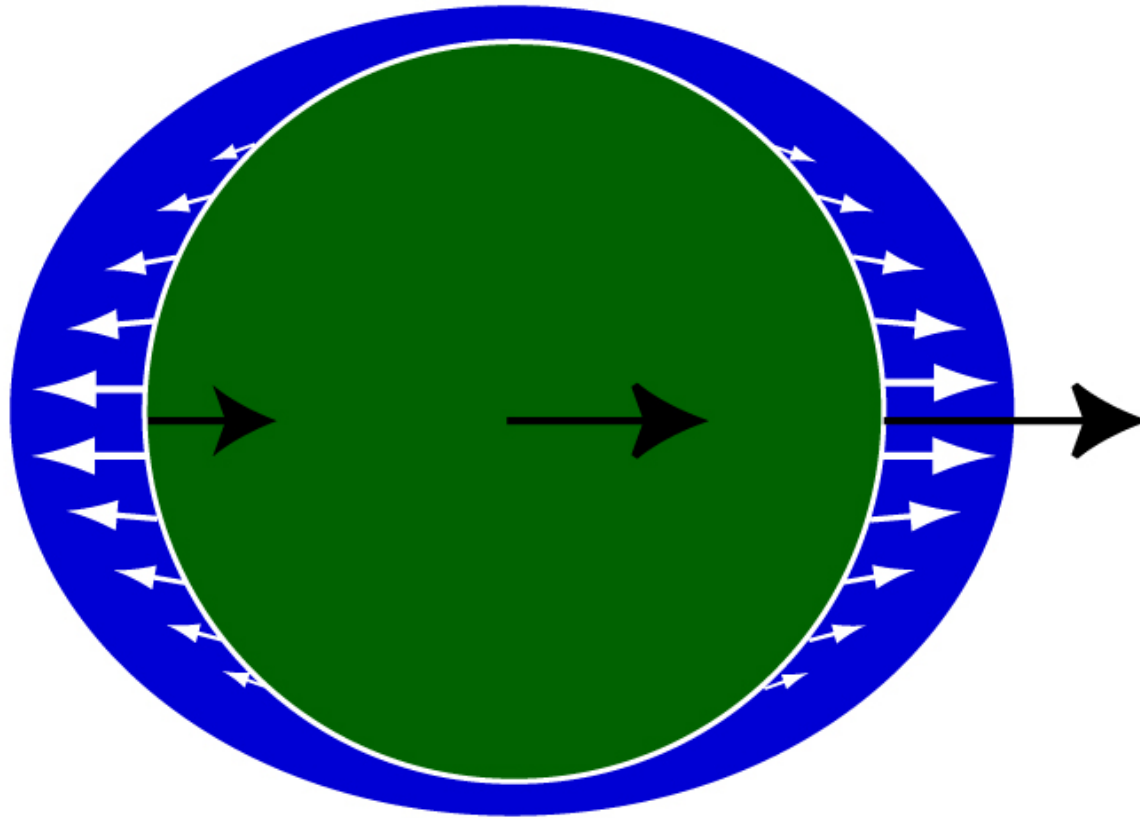
The rings could have been much more massive in the past.

TECHNICAL INTRODUCTION

1) Kepler's Law

2) Roche's radius

TIDES



Black arrows: Gravitational force due to Moon.
White arrows: Net differential force relative to centre of the Earth - the tide-raising force.

TIDES

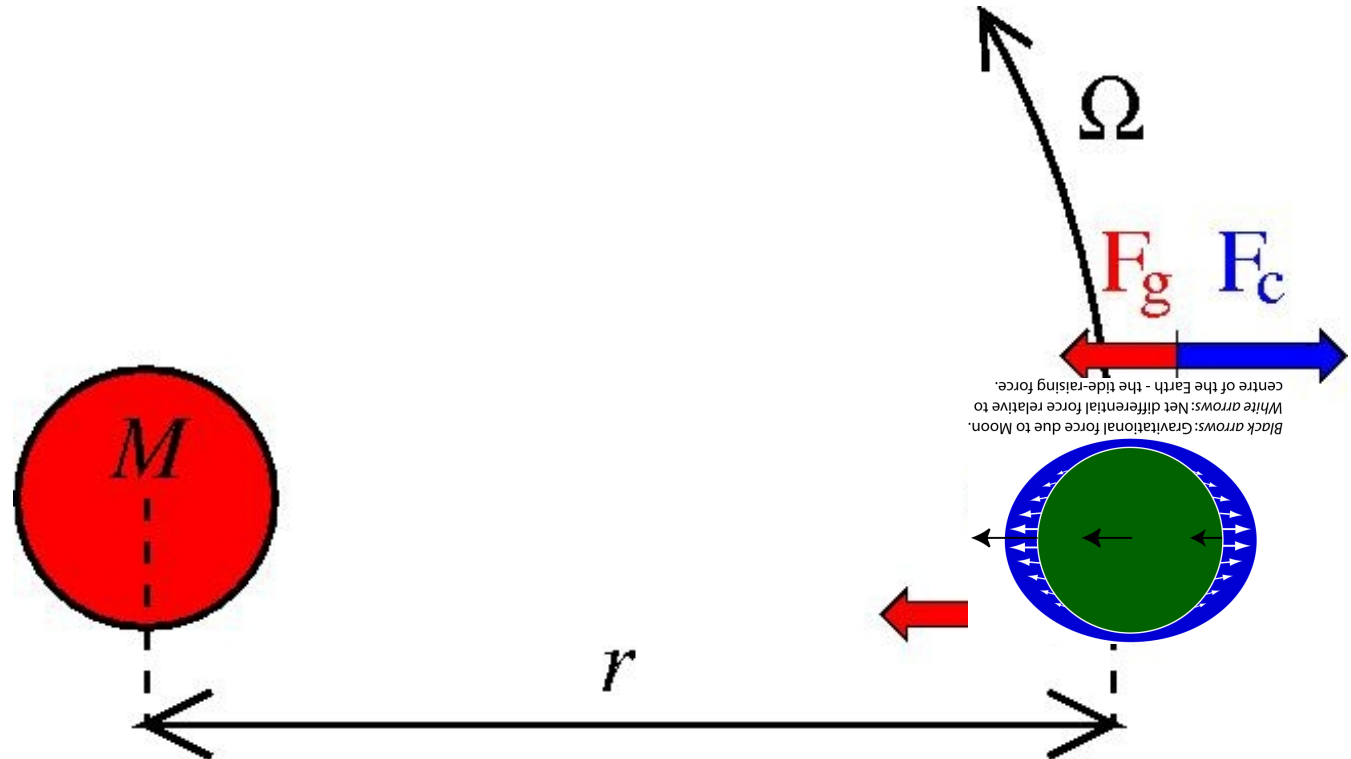
Reminder : Tidal forces (per mass unit) :

$$\Omega = (GM/r^3)^{1/2}$$

$$F_g = GM / (r \pm a)^2$$

$$F_c = \Omega^2(r \pm a)$$

$$F_t = 3\Omega^2 a$$



Roche Radius

Self-gravity force of the two spheres (per mass unit) :

$$F_{sg} = G^*(4/3)\pi\rho a^3 / (2a)^2$$

Condition for stability of the aggregate : $F_{sg} > F_t$,

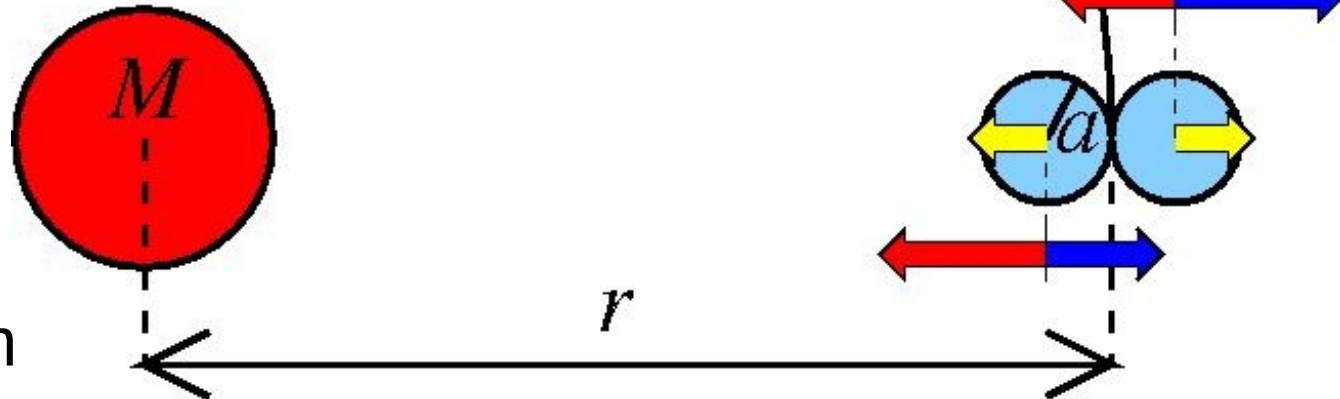
or : $r > (9M/\pi\rho)^{1/3} = r_{\text{Roche}}$

Application:

$$M = M_{\text{Saturn}},$$

$$\rho = 600 \text{ kg.m}^{-3}$$

$$r_{\text{Roche}} = 1,4 \cdot 10^8 \text{ m}$$



Application :

Saturn's rings
extend to
136 000 km.

Composition:

10cm – 10m
blocks, of 90%
water ice !

→ Tidal disk.

Movie:
Hanno REIN.



TECHNICAL INTRODUCTION

SUMMARY:

1) The rings can't agglomerate because they are inside the Roche radius.

2) But they spread...
(and were much more massive in the past)

3) Satellites migrate outwards, repelled by the rings.

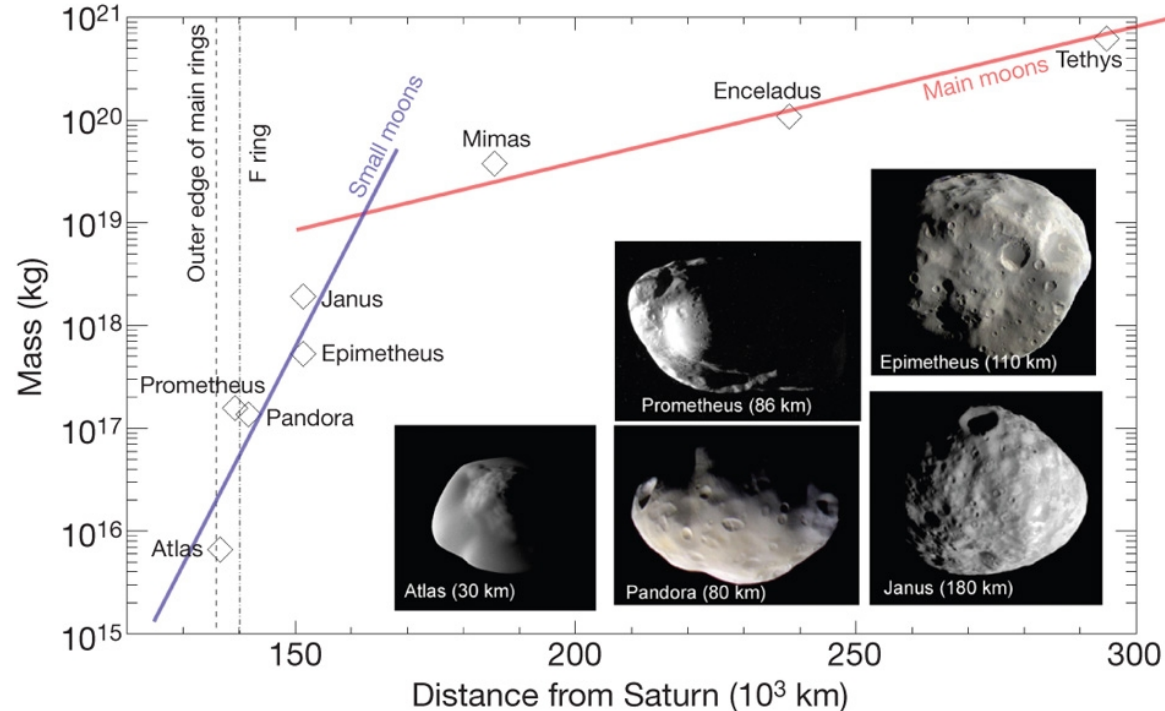
In particular, Janus should have been inside the rings 100 Myrs ago !

Satellites children of the rings

1) Saturn's small moons

Just outside Saturn's rings, there is a handful of small satellites, with surprising properties :

- underdense ($\sim 600 \text{ kg.m}^{-3}$)
- same spectrum as the rings
- dynamically young
- young surfaces



Satellites children of the rings

1) Saturn's small moons

But... we have seen that the rings spread...

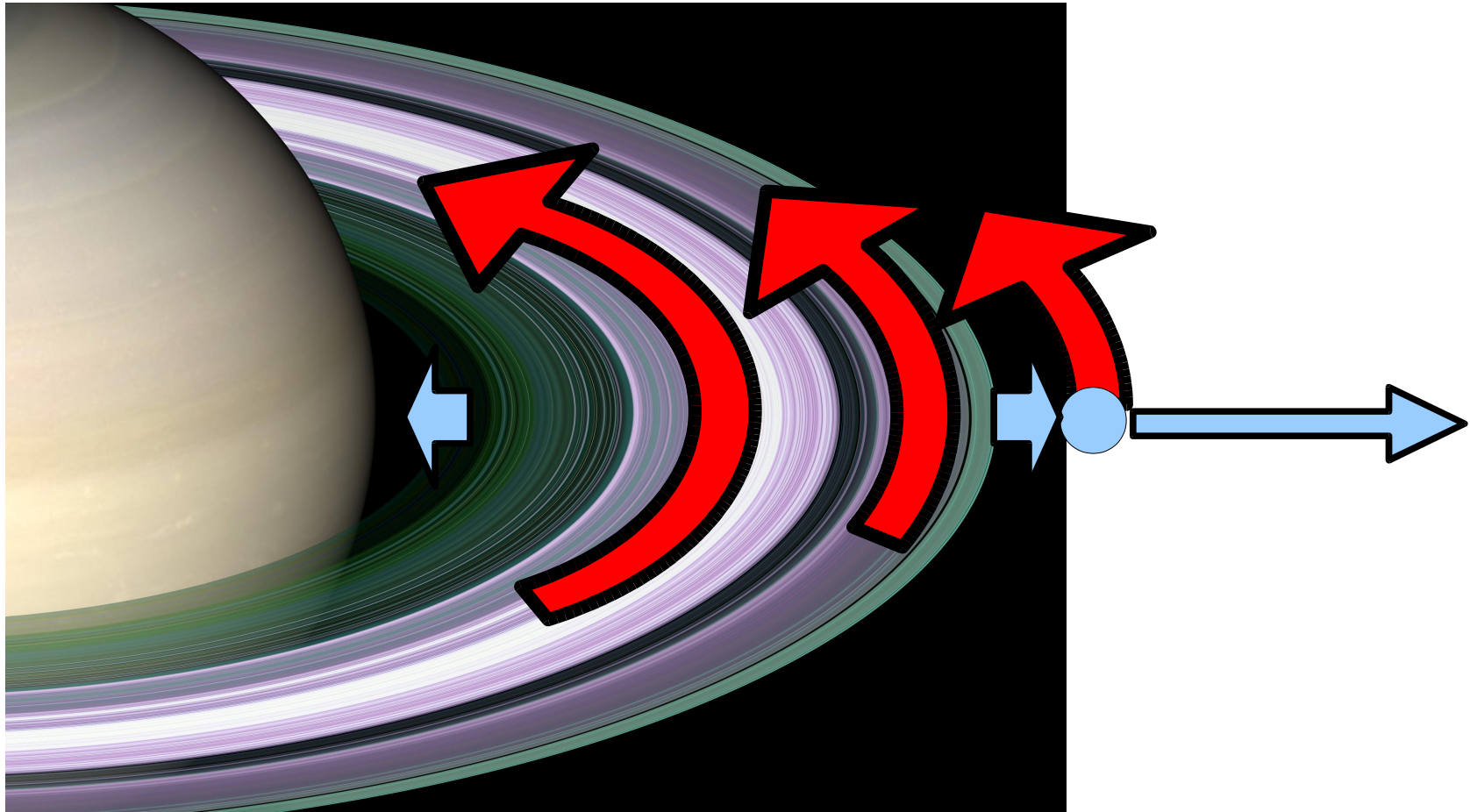
On the inside, the ice falls into Saturn.

And on the outside ?

After crossing the Roche limit,
the ice boulders agglomerate, accrete, coalesce,
and form new small satellites !

Satellites children of the rings

1) Saturn's small moons



Satellites children of the rings

1) Saturn's small moons

We made numerical simulations of this process, with viscous spreading of the rings, and satellite formation beyond r_{Roche} (Charnoz, Salmon, Crida, 2010).

It works !

In ~100 Myrs, reproduction of the 5 smallest moons from rings like today's.

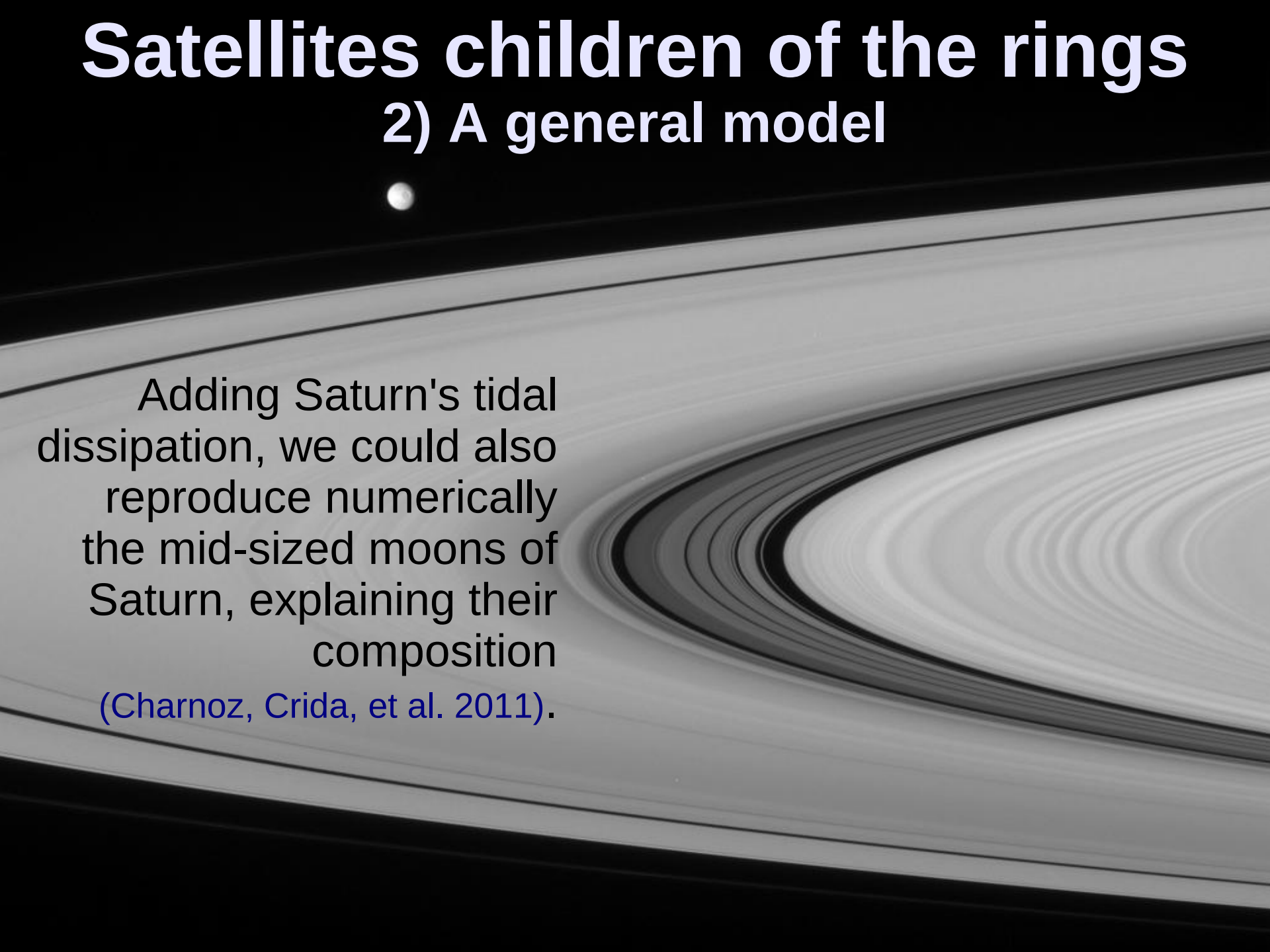
Satellites children of the rings

1) Saturn's small moons



Satellites children of the rings

2) A general model

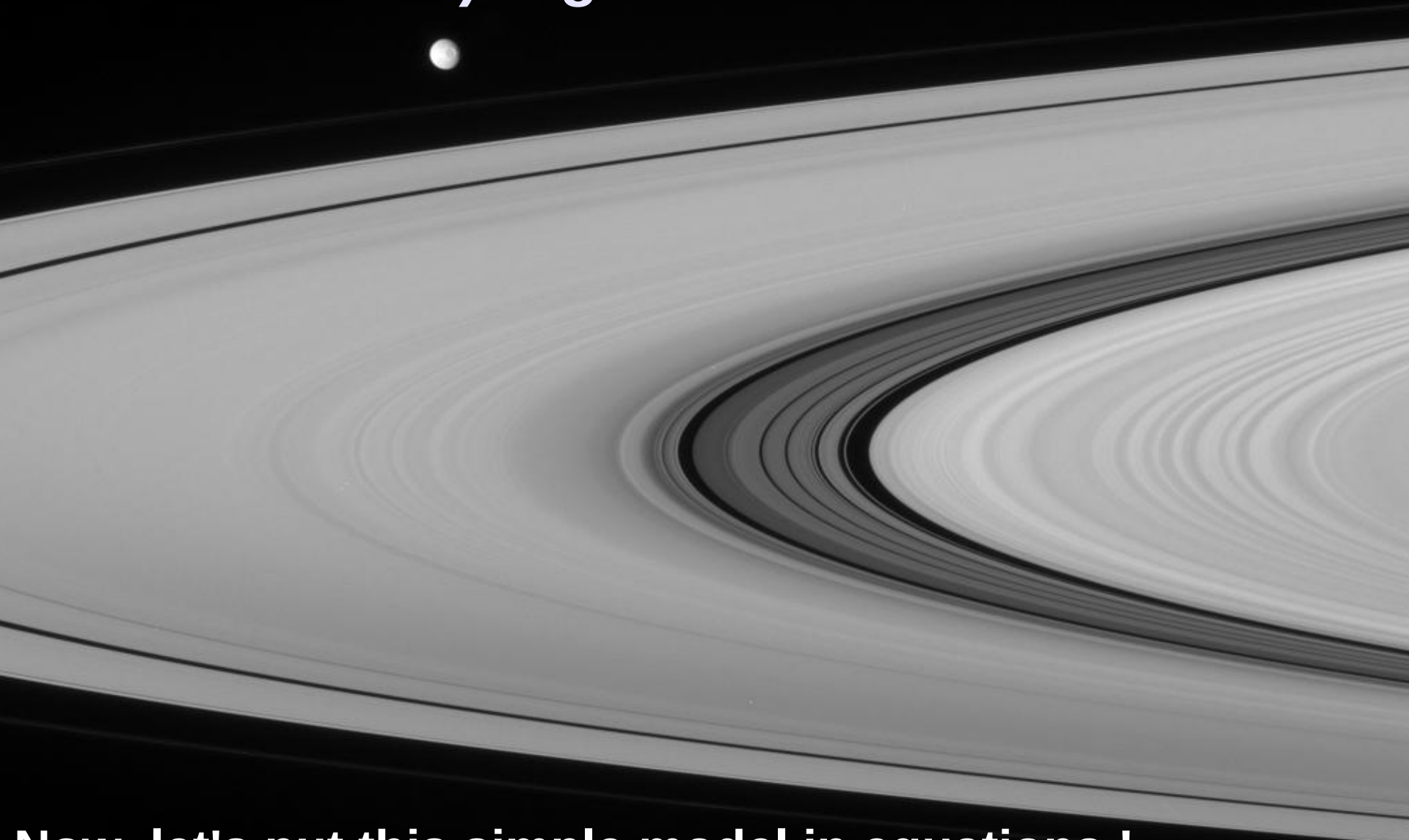


Adding Saturn's tidal dissipation, we could also reproduce numerically the mid-sized moons of Saturn, explaining their composition

(Charnoz, Crida, et al. 2011).

Satellites children of the rings

2) A general model

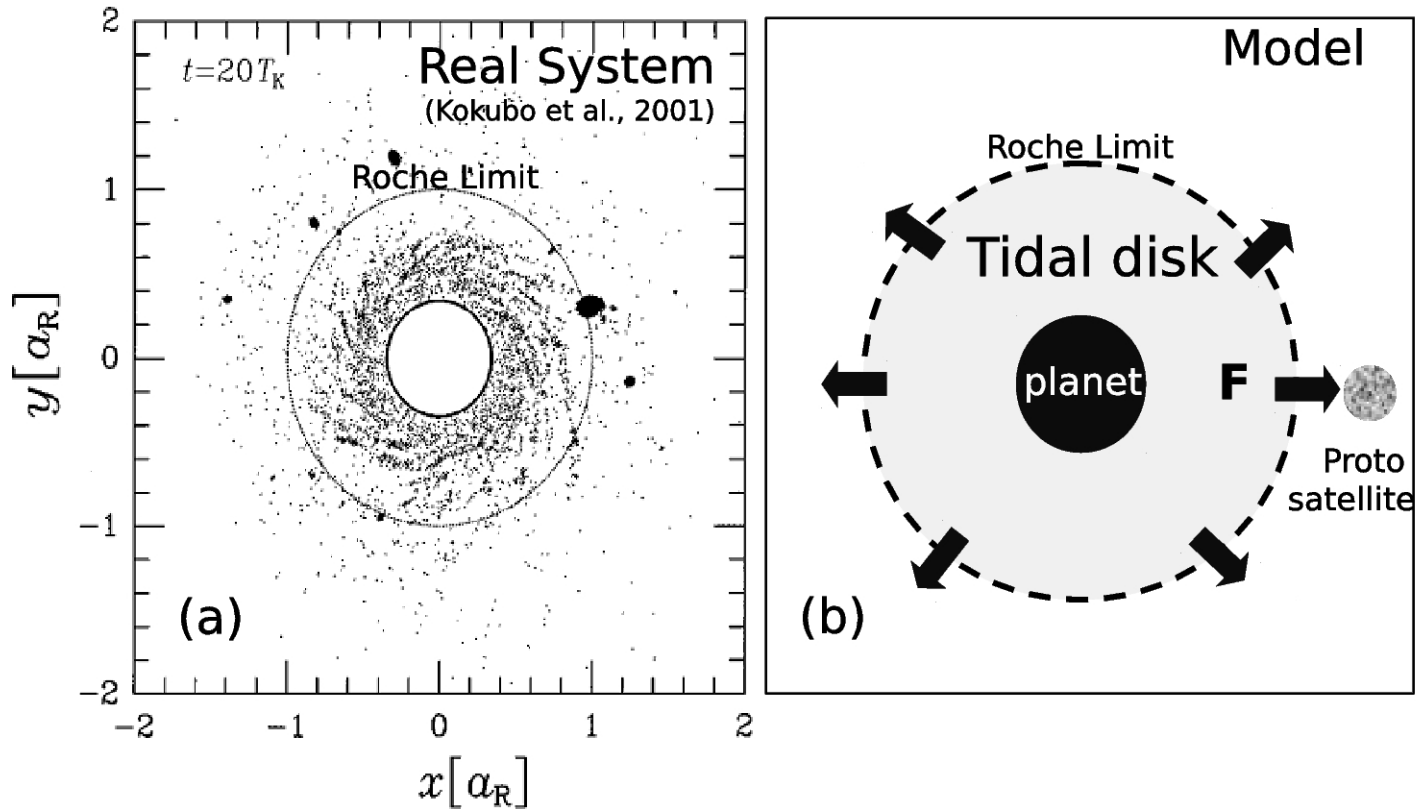


Now, let's put this simple model in equations !

Spreading of a tidal disk

Inside the Roche radius r_R , there is a « tidal disk », that spreads with a mass flow F (assumed constant).

Beyond r_R , new satellite(s) form...



Notations

1D model.

Be T_R the orbital period at r_R , and

$\tau_{\text{disk}} = M_{\text{disk}} / FT_R$, the normalized life-time of the disk.

The disk spreads with a viscous time $t_v = r_R^2 / \nu$.

Using Daisaka et al. (2001)'s prescription for ν ,
we find : $\tau_{\text{disk}} = t_v / T_R = 0.0425 D^{-2}$ where $D = M_{\text{disk}} / M_p$.

Continuous regime

Say 1 satellite forms. Its mass is : $M = F t$ (1)

It feels a torque from the tidal disk : $\Gamma = \frac{8}{27} \left(\frac{M}{M_p} \right)^2 \Sigma r^4 \Omega^2 \Delta^{-3}$

where $\Delta = (r - r_R) / r_R$ (Lin & Papaloizou 1979).

→ Migration rate :

$$\frac{d \Delta}{d t} = \frac{32}{27} q D T_R^{-1} \Delta^{-3} \quad (2)$$

where $q = M / M_p$.

Solution of (1) & (2) :

$$q = \left(\frac{\sqrt{3}}{2} \right)^3 \tau_{disk}^{-1/2} \Delta^2 \quad (3)$$

We call this the *continuous regime* .

Continuous regime

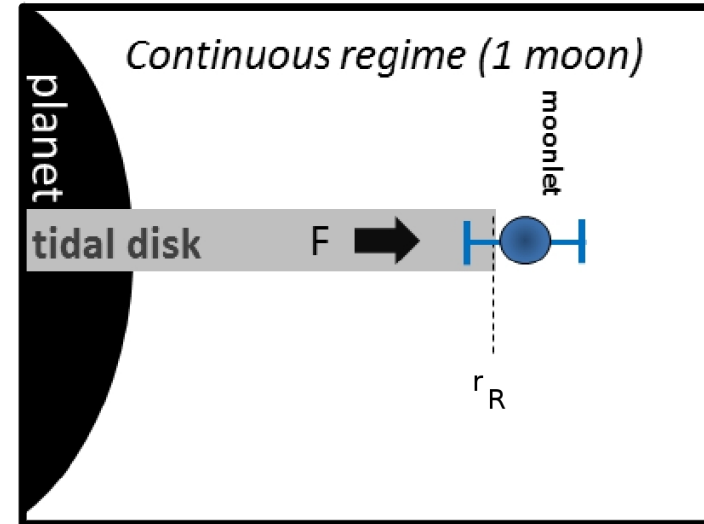
This holds as long as the satellite captures immediately what comes through r_R .

That is, as long as $(r-r_R) < 2 r_{\text{Hill}}$,
or $\Delta < 2 (q/3)^{1/3}$.

Input into Eq.(3), this gives a condition of validity for the continuous regime :

$$\Delta < \Delta_c = \sqrt{\frac{3}{\tau_{\text{disk}}}} = \sim 8.4 D$$

$$q < q_c = \frac{3^{5/2}}{2^3} \tau_{\text{disk}}^{-3/2} = \sim 222 D^3$$



Duration of the continuous regime: $10 T_R$.

Applications

Using the relation between τ_{disk} and D , one finds :

1) Earth's Moon forming disk :

$$q_c = \sim \text{mass of the Moon !}$$

=> Agreement with N-body numerical simulations.

2) Charon never left the continuous regime.

3) Saturn's rings now :

$$q_c = \sim 10^{-18} .$$

=> Only very small objects are formed.

Pyramidal regime

Satellites of mass q_c are produced at Δ_c every q_c / F .

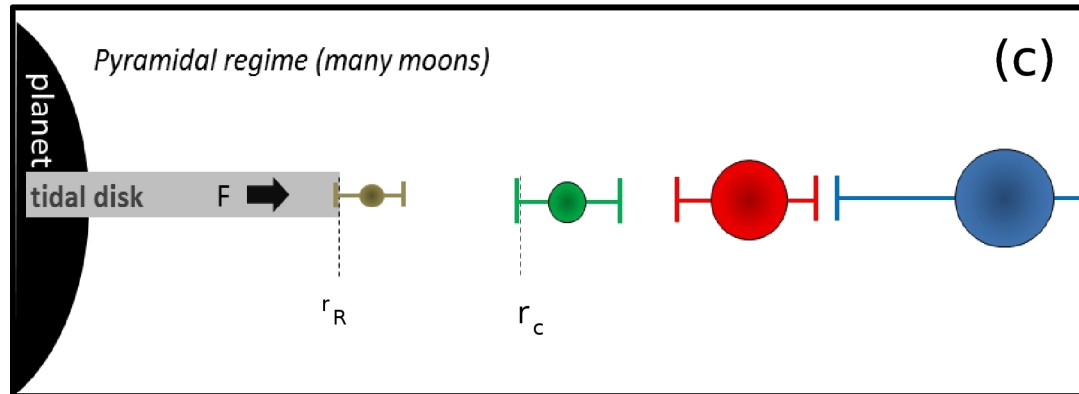
Then, many satellites of constant mass migrate outwards, at decreasing rates. They approach each other.

If their distance decreases below 2 mutual Hill radii, they merge.

This leads to the formation of satellites of masses $2q_c$, every $2q_c / F$. They migrate away and merge further...

And so on, hierarchically...

We call this *the pyramidal regime*.



Pyramidal regime

- Using Eq.(2), we show that in the pyramidal regime, while the mass is doubled, Δ is multiplied by $2^{5/9}$.

Thus, $q \propto \Delta^{9/5}$.

In addition, the number density of satellites should be proportionnal to $1/\Delta$, explaining the pile-up.

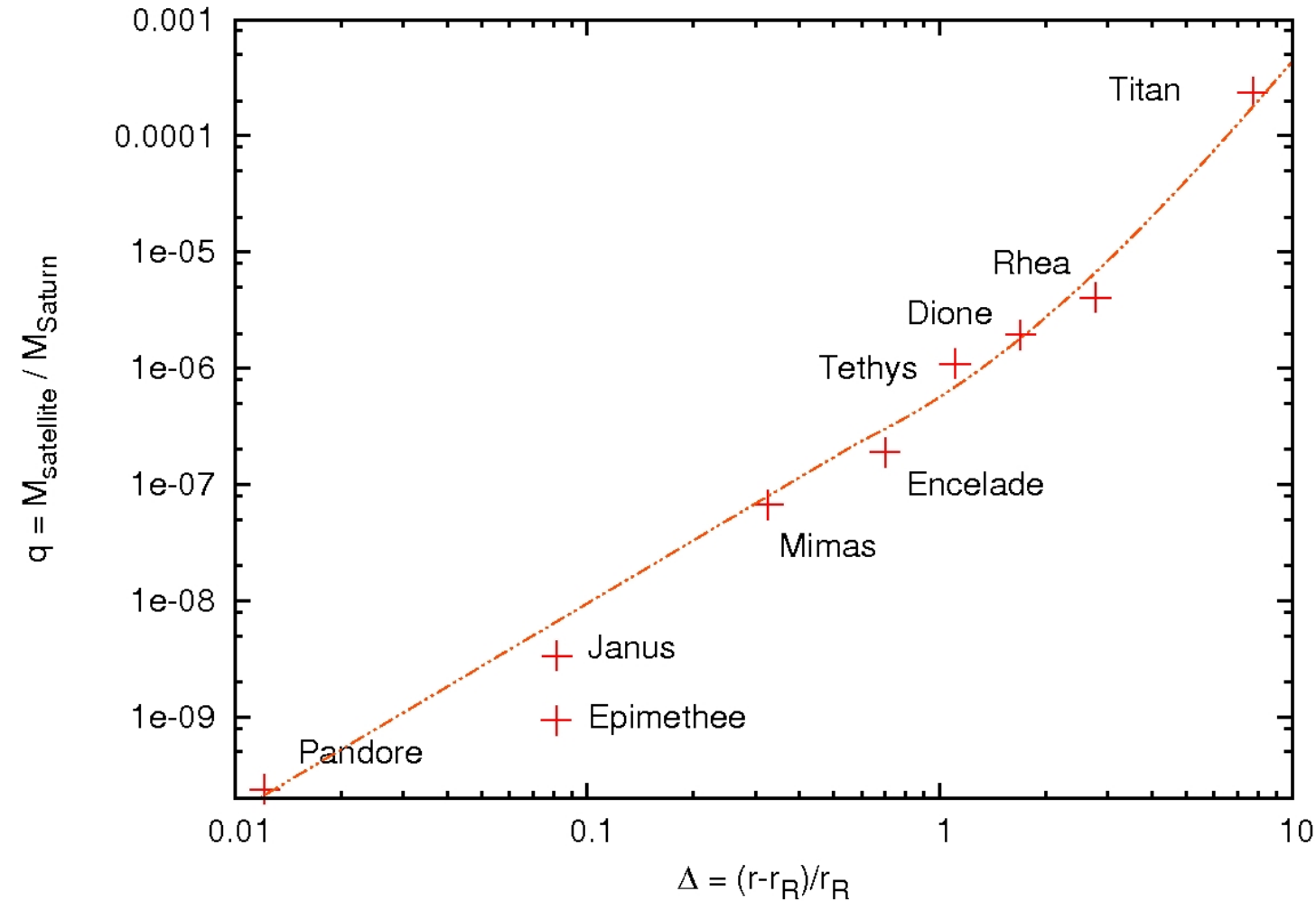
- Beyond the 2:1 Lindblad resonance with r_R ($\Delta=0.58$), Eq.(2) doesn't apply. Migration is driven by planetary tides:

$$\frac{dr}{dt} = \frac{3 k_{2p} M \sqrt{G} R_p^5}{Q_p \sqrt{M_p} r^{11/2}} \quad (4)$$

Using Eq.(4), we find $q \propto r^{3.9}$.

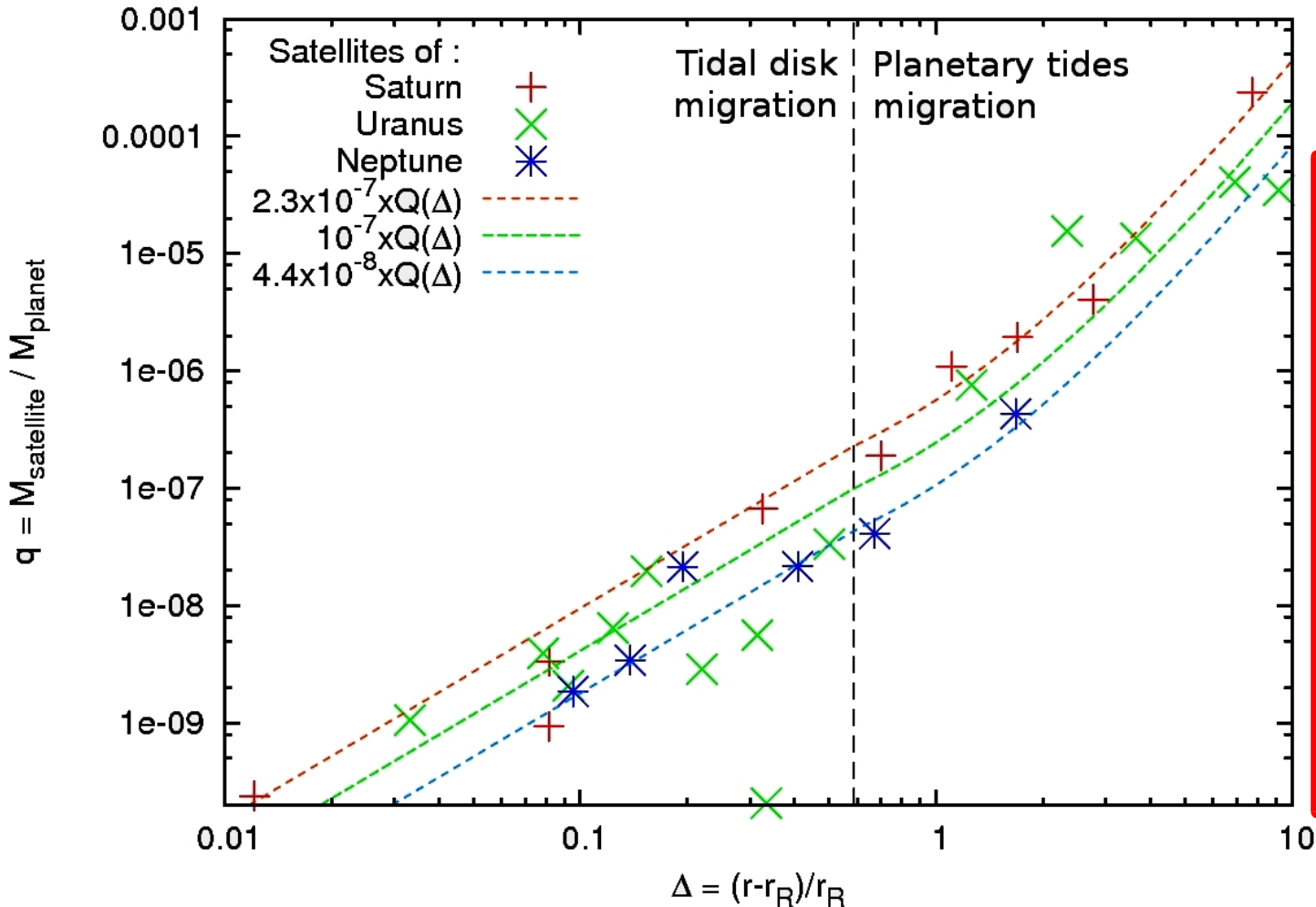
Pyramidal regime

The result spectacularly matches the distribution of the Saturnian system !



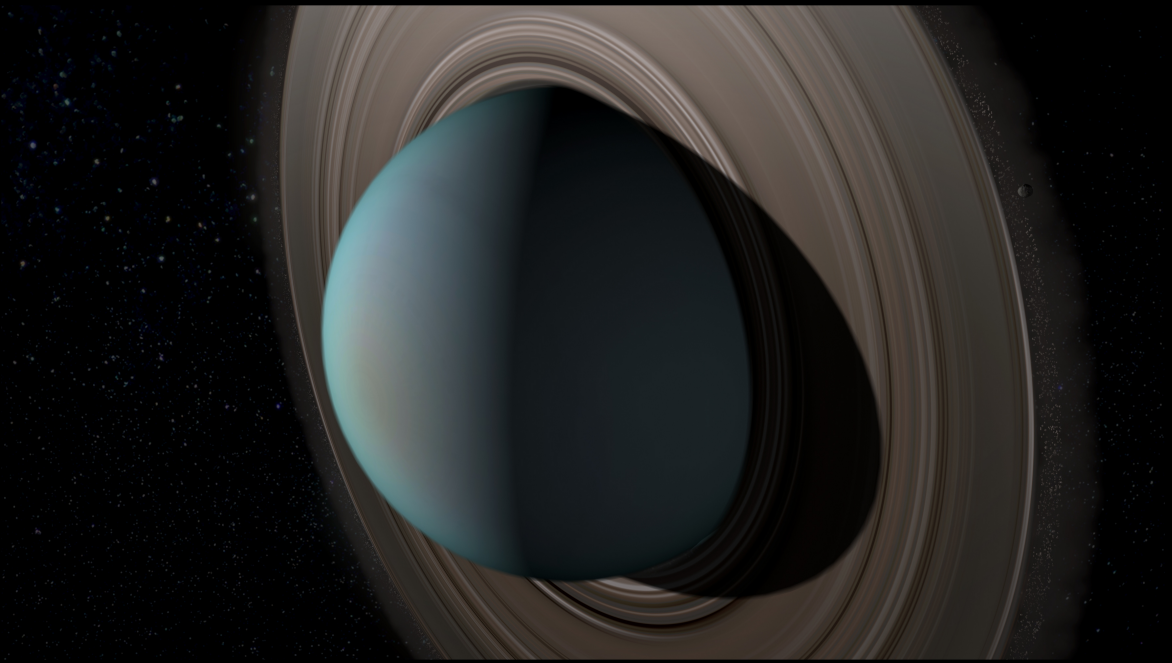
Pyramidal regime

The result spectacularly matches the distribution of the Saturnian, Uranian, and Neptunian systems !



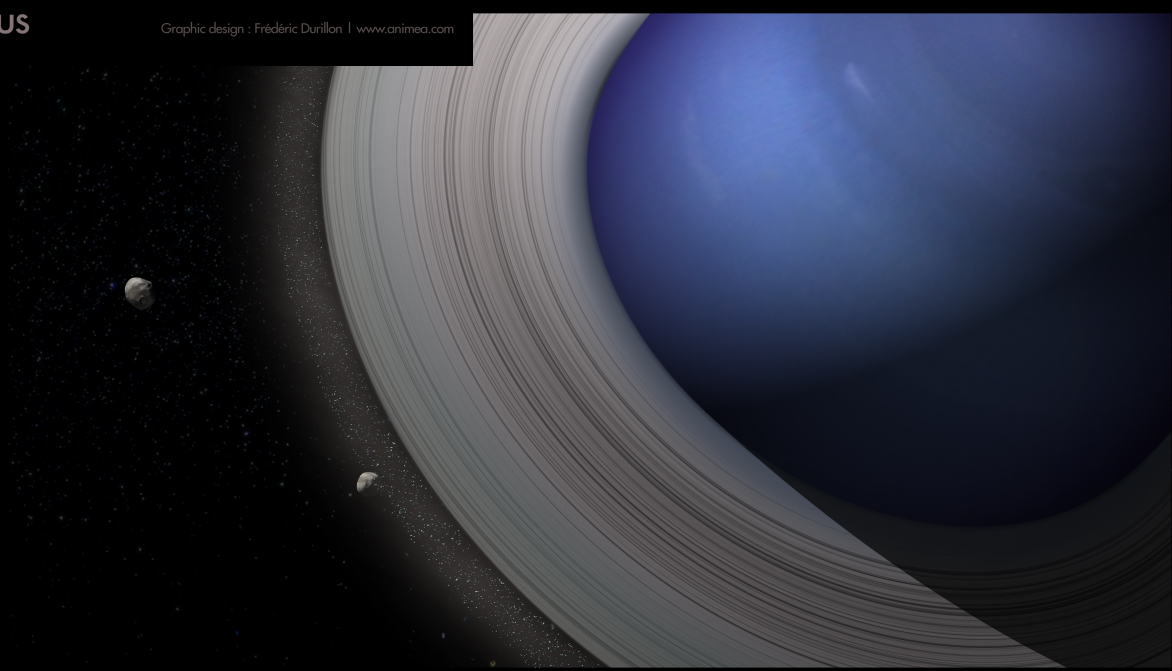
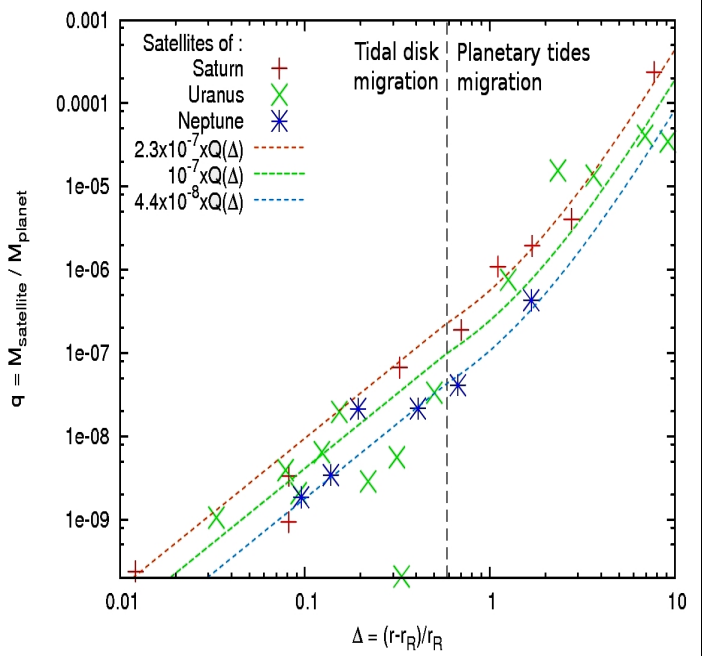
I claim that Uranus and Neptune had massive rings, from which their regular satellites were born.

I claim that Uranus and Neptune had massive rings, from which their regular satellites were born.



Uranus

Graphic design : Frédéric Durillon | www.animeca.com



Neptune

Graphic design : Frédéric Durillon | www.animeca.com



Summary

1) Continuous regime:

1 moon grows

$$q \propto \Delta^2$$

until Δ_c or q_c .

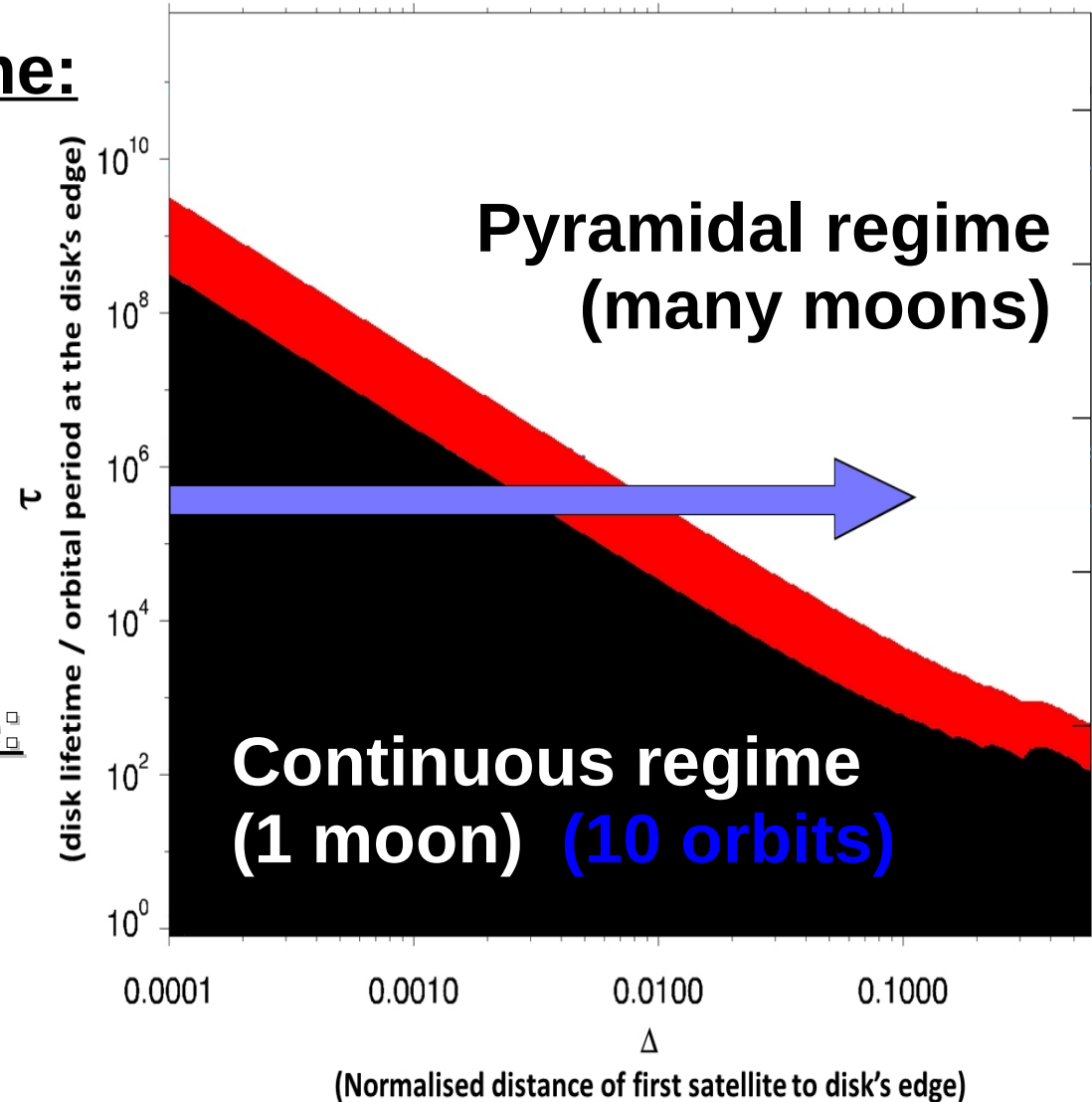
2) Discrete regime:

transition, 2 moons

3) Pyramidal regime:

Many moons

$$q \propto \Delta^{9/5} \text{ or } r^{3.8} .$$



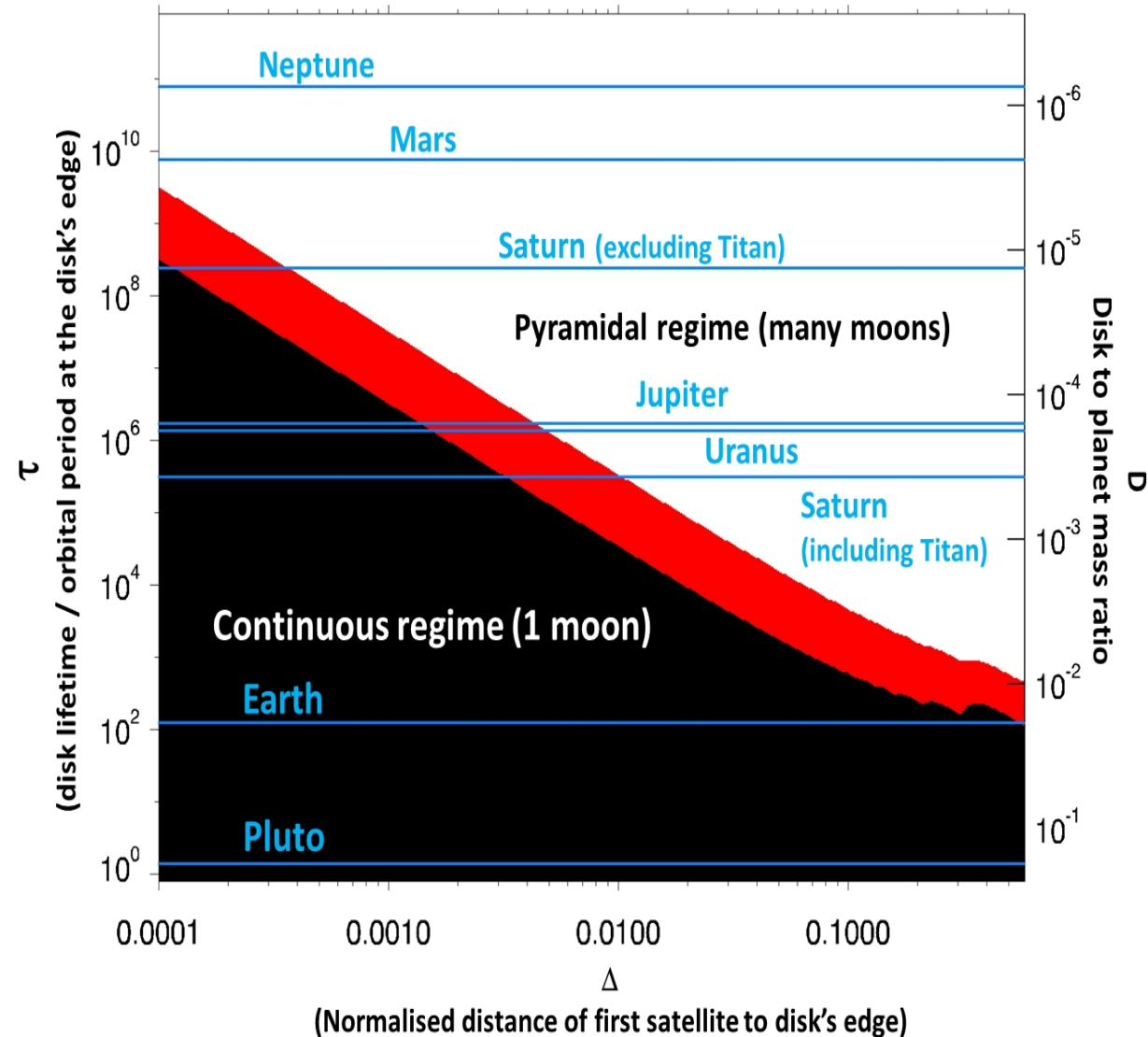
Summary

Application :

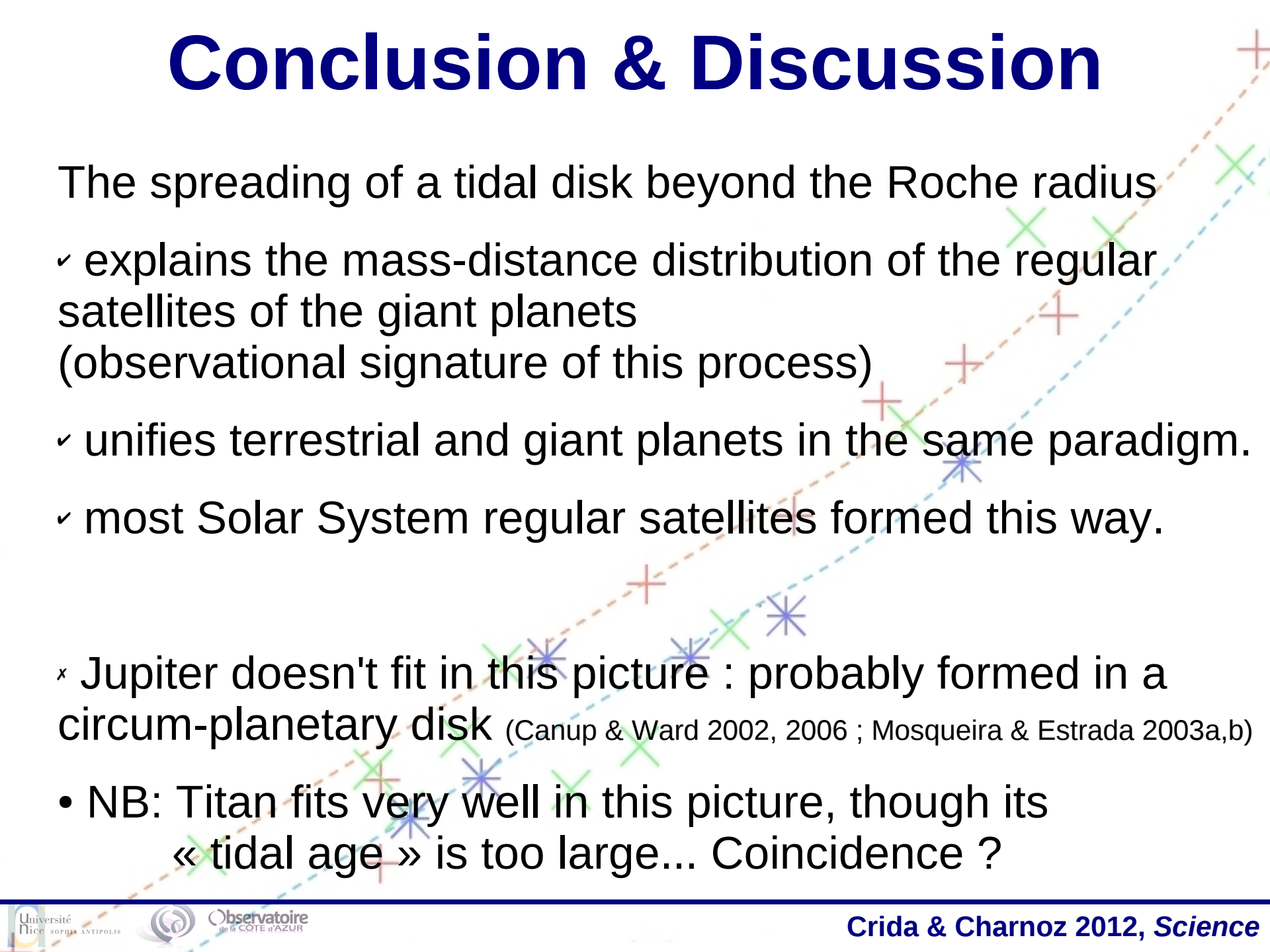
Take $M_{\text{disk}} = 1.5 \times$
total mass of present
satellite system.

Giant planets :
dominated by the
pyramidal regime,

Earth and Pluto :
1 large satellite.



Conclusion & Discussion

- The spreading of a tidal disk beyond the Roche radius
- ✓ explains the mass-distance distribution of the regular satellites of the giant planets (observational signature of this process)
 - ✓ unifies terrestrial and giant planets in the same paradigm.
 - ✓ most Solar System regular satellites formed this way.
- × Jupiter doesn't fit in this picture : probably formed in a circum-planetary disk (Canup & Ward 2002, 2006 ; Mosqueira & Estrada 2003a,b)
- NB: Titan fits very well in this picture, though its « tidal age » is too large... Coincidence ?
- 

Thanks !

Aurélien CRIDA

& Sébastien CHARNOZ



Observatoire
de la CÔTE d'AZUR

ORIGIN of the SATELLITES ?

As for our Moon, the Earth didn't have a Circum-Planetary Disk, but a giant impact with a Mars-sized body ejected enough material (and angular momentum) around the Earth to form the Moon

(Canup & Asphaug 2001)

